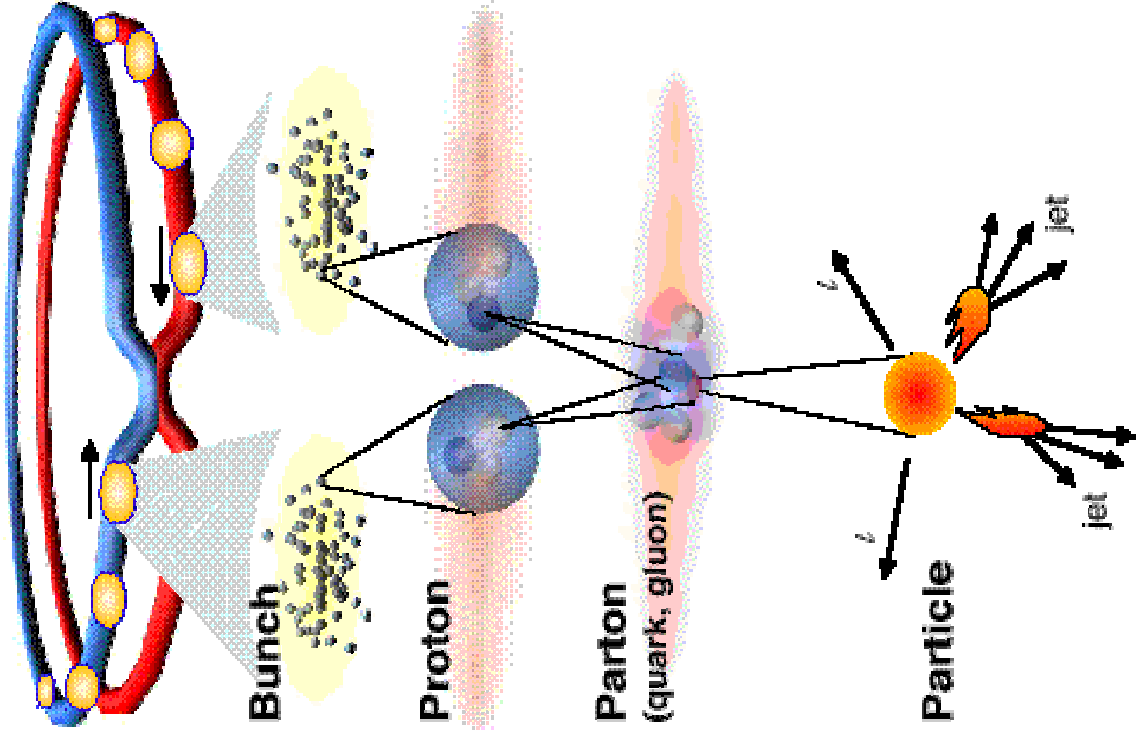


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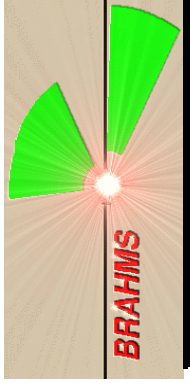
# Particle ratios (and more) from proton + proton collisions



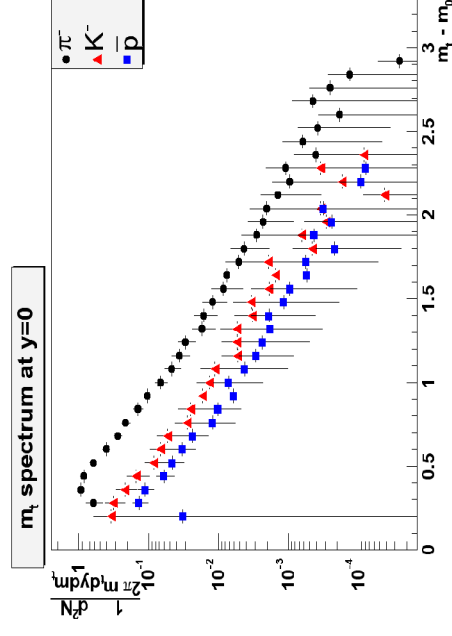
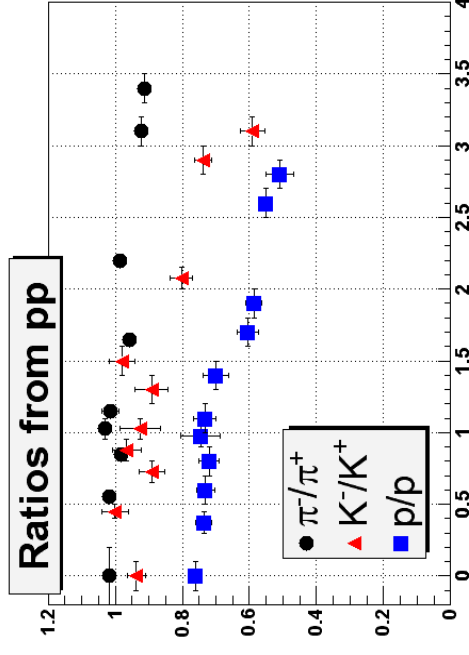
*p+p* ratios  
Bjørn H. Samset

# Contents

- Data selection
- Analysis framework
- PID
- Main analysis
- Results: Ratios vs.  $y$  and  $p_t$   $++$
- Problems and caveats
- Towards yields: MRS acceptances and prelim. spectra
- Discussion: FS acceptances?!?

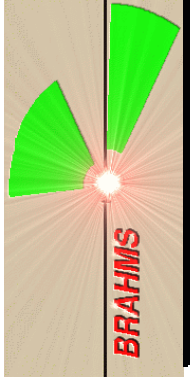


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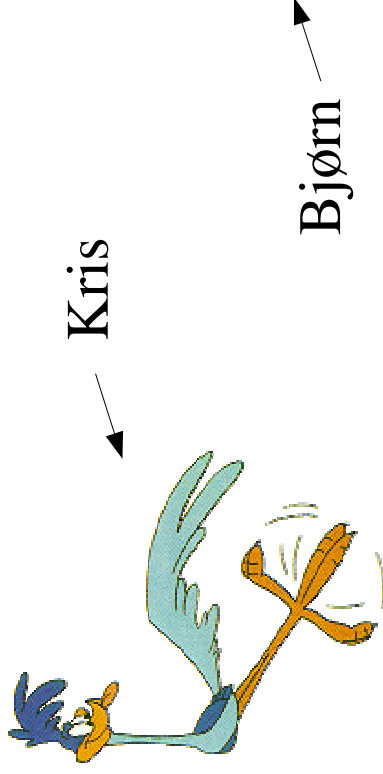
*p+p ratios*  
Bjørn H. Samset

# Data, local & global tracking



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- Data from 2001 only (so far...)
- Local tracking, global tracking, calibrations etc. by Kris Hagel



- All the data have recently been (re-re-)reanalyzed, taking into account new calibrations and a couple of bug fixes (KH, FV)
- Event sample: 11.5M events...

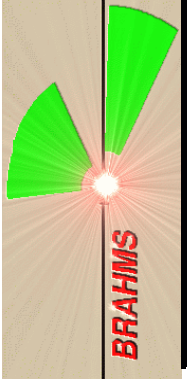


*p+p ratios*  
Bjørn H. Samset

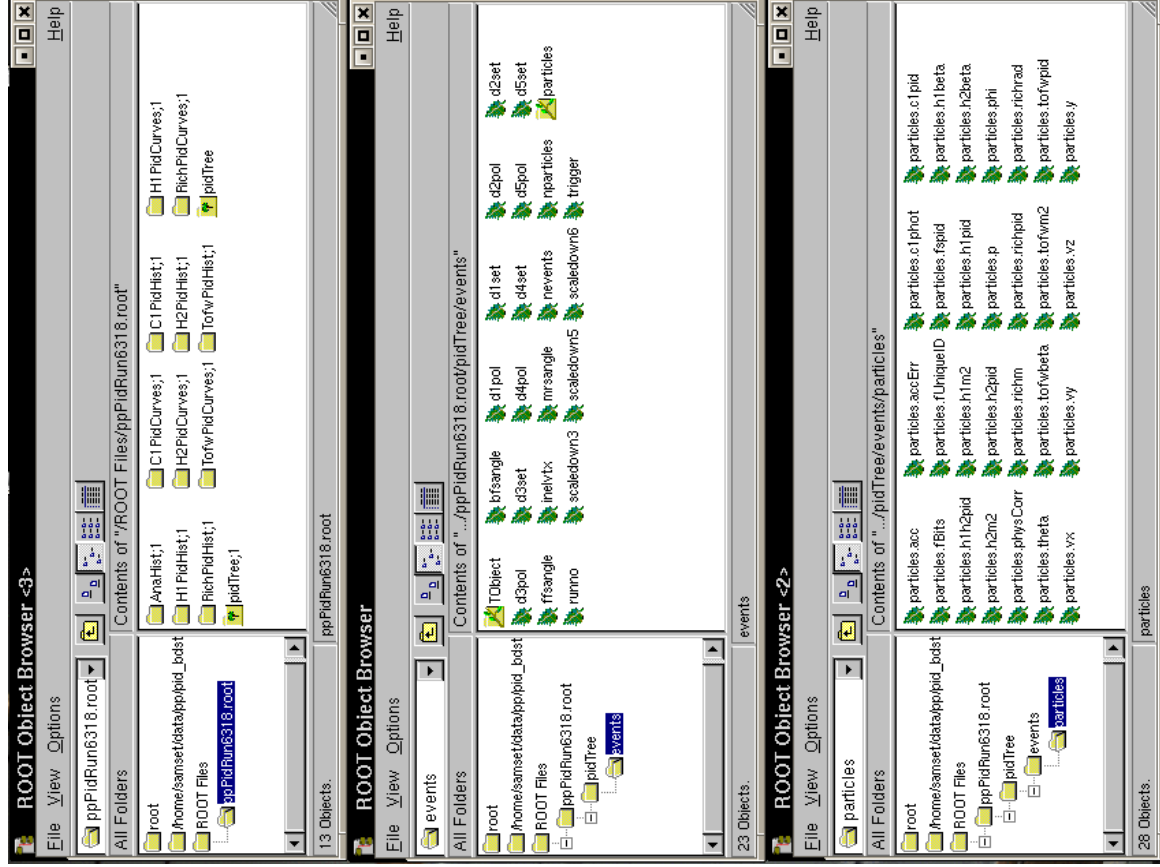
# Analysis framework:

# FLAP

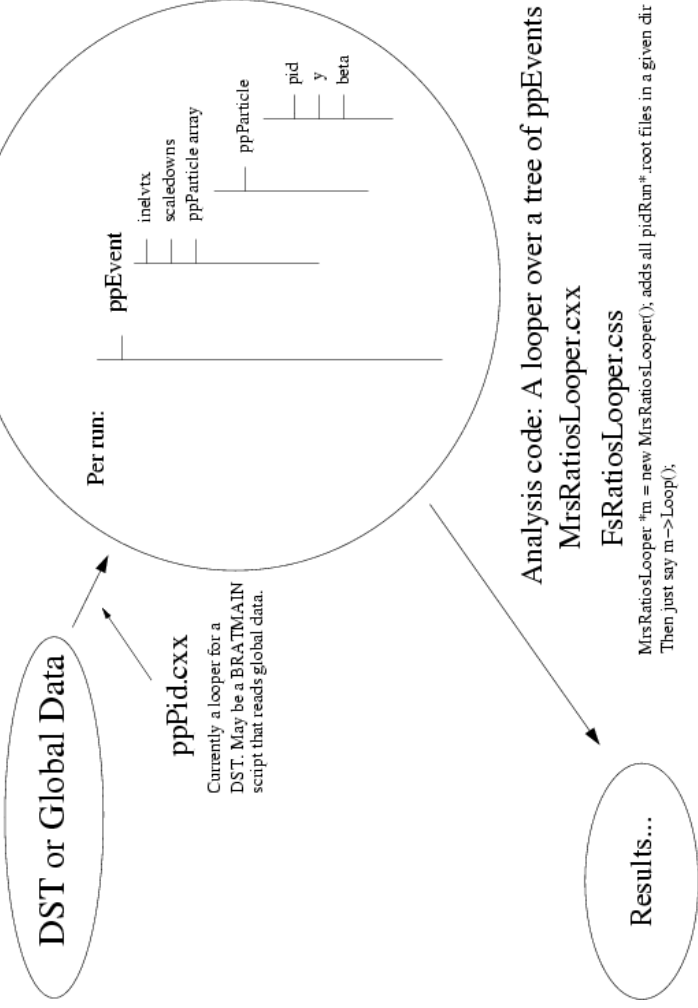
(Final Looping Analysis and Pid)



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pp Analysis Flowchart

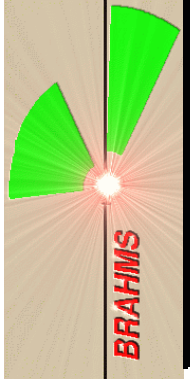


- Analysis is quick (~1hr for full 2001 dataset)
- Files are small (full dataset is 350Mb)

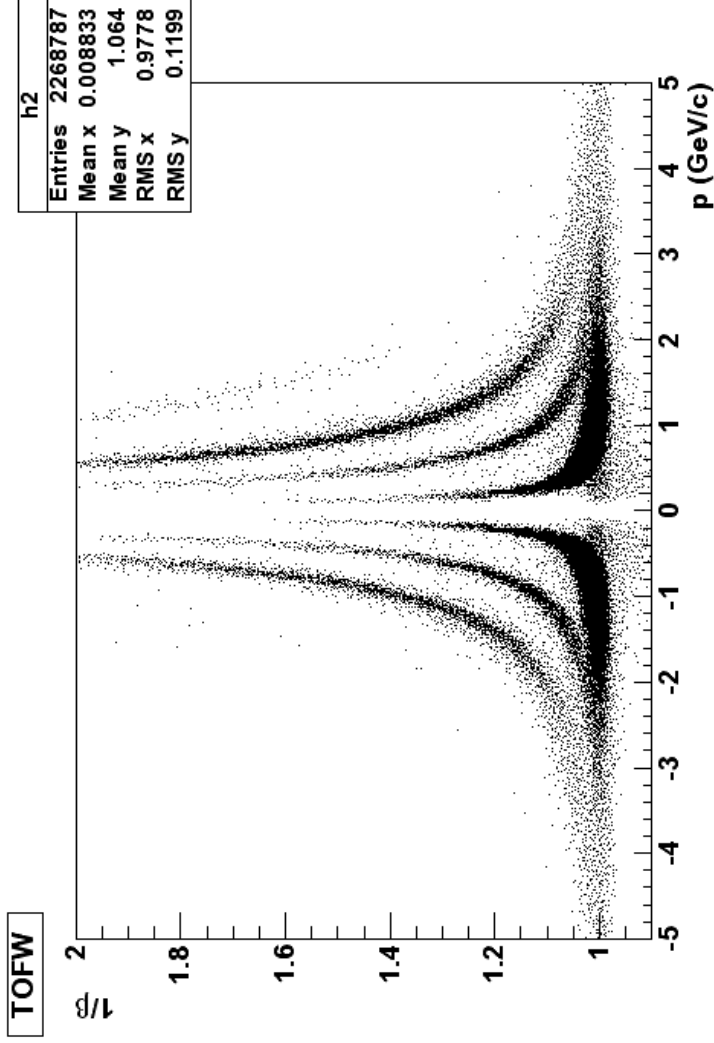


*p+p* ratios  
Bjørn H. Samset

# PID - TOFW



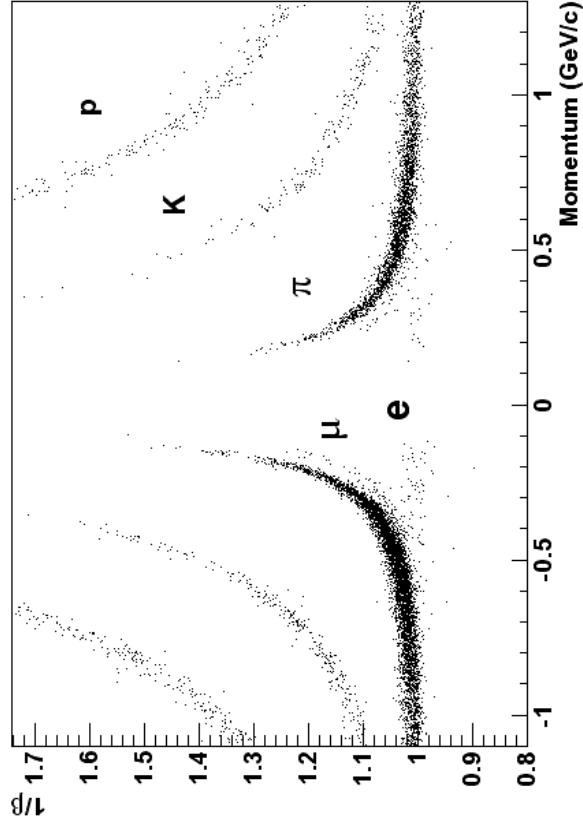
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All MRS data.

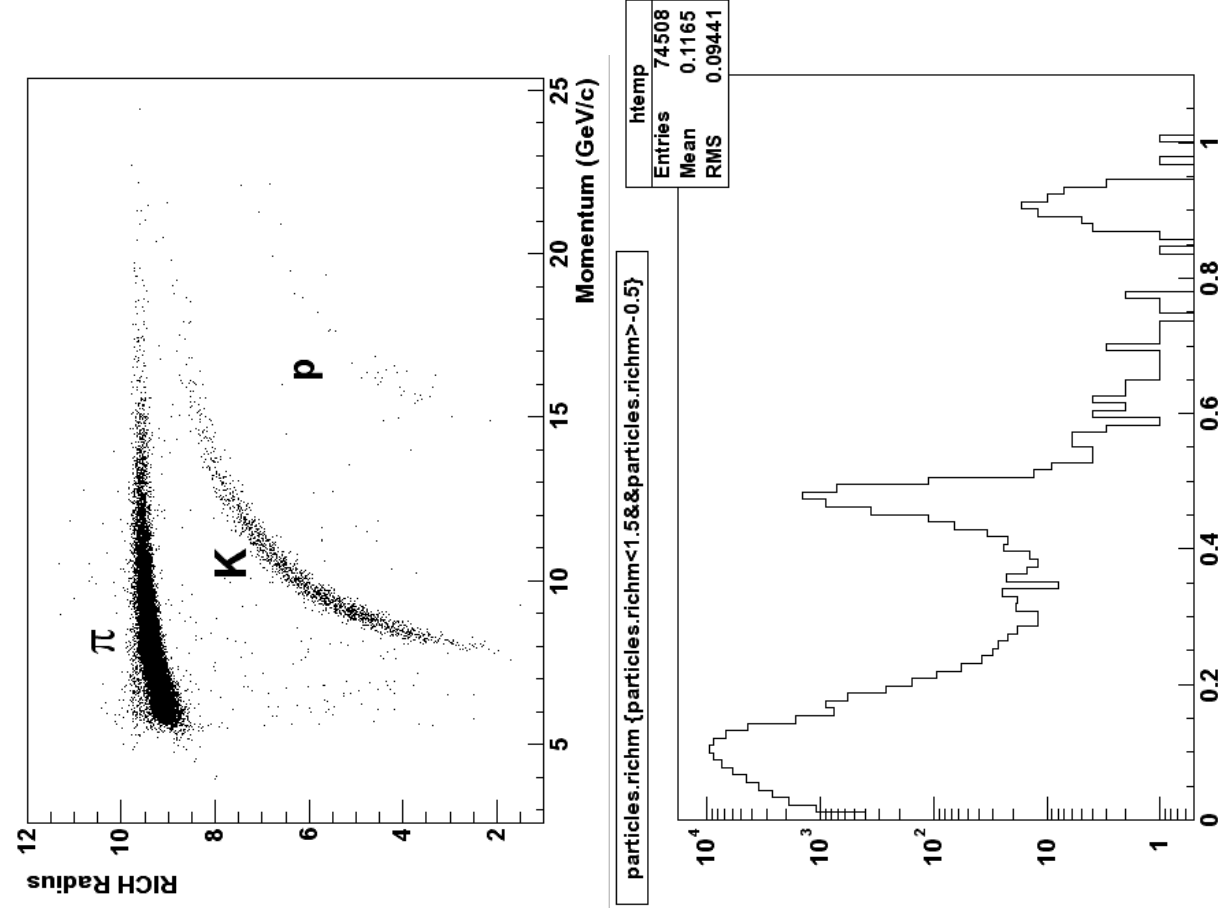
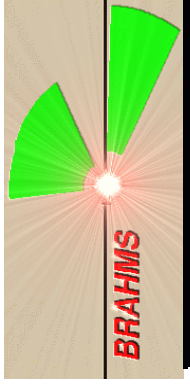
- p/K up to  $\sim 3$  GeV/c
- K/ $\pi$  up to  $\sim 1.5$  GeV/c

Single 90deg run. Good, clean  
PID for p, K,  $\pi$ ,  $\mu$  and e...

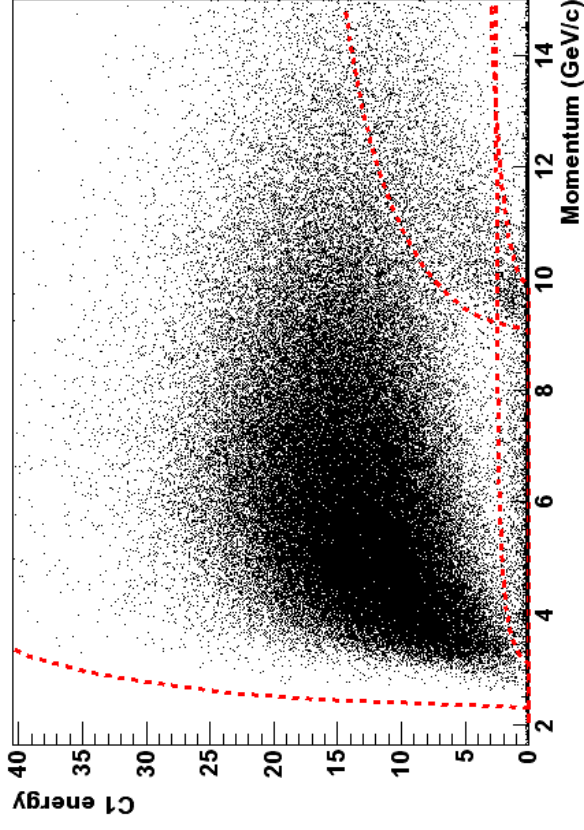


*p+p ratios*  
Bjørn H. Samset

# PID - Cherenkovs



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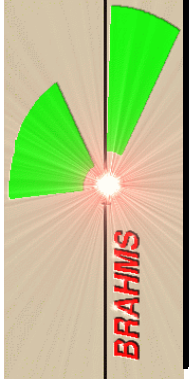
FS PID sequence:

- C1 (pions, kaons only)
- RICH
- H1 with pions vetoed
- H2 with pions vetoed

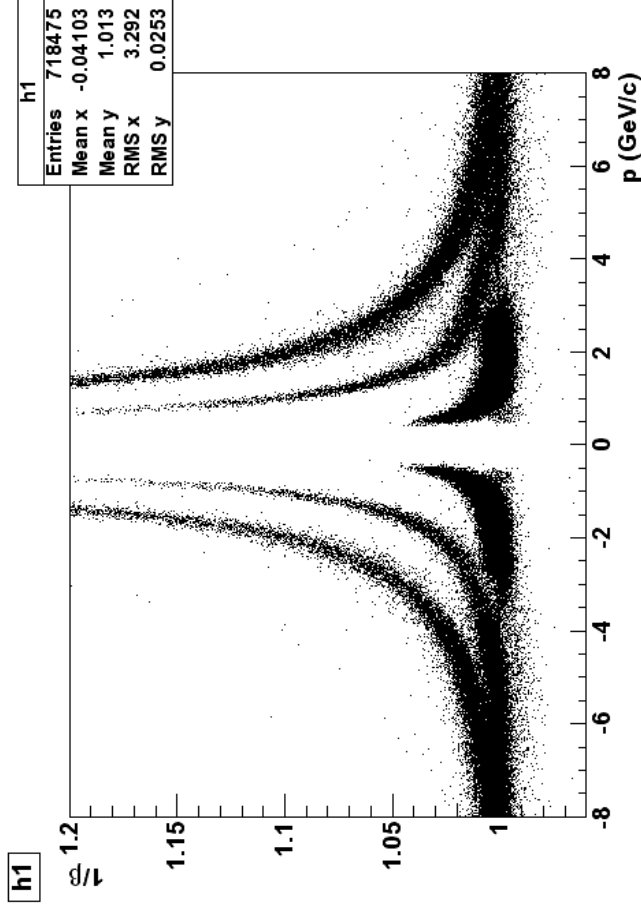
*p+p ratios*  
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# PID - H1 and H2

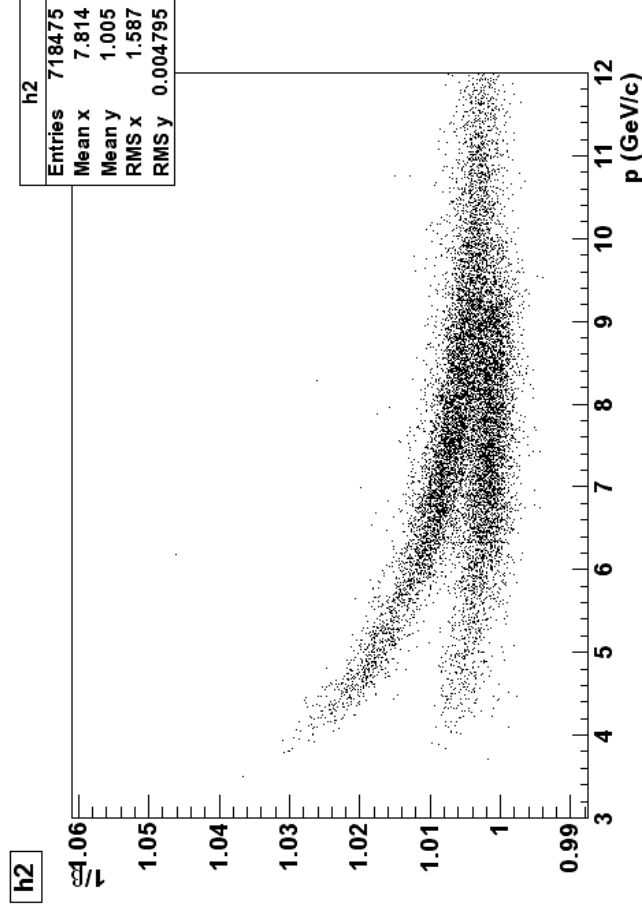


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p/K up to  $\sim 4.5$  GeV/c  
pi up to  $\sim 3.0$  GeV/c

**NB: Momentum!**  
p is average of d2, d3 and d4 mom.



p/K up to  $\sim 6.5$  GeV/c? 7 GeV/c?  
No pions (all vetoed by C1/RICH)



*p+p ratios*  
Bjørn H. Samset

# Run selection

Requirements to accept a run:

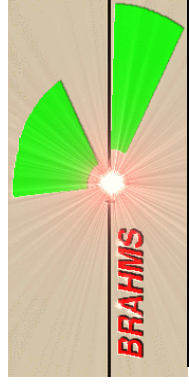
- PID looks OK for all detectors
- Approx. constant particles/trigger for all runs in same setting, checked for indiv. detectors
- FS: We have data for both A and B pol. at roughly the same angle and magn. field.

FS runs used:

6256 6303 6418 6435 6449  
6257 6400 6427 6442 6473  
6277 6401 6428 6445 6475  
6278 6405 6429 6446 6488  
6279 6410 6431 6447 6493  
6302 6411 6434 6448 6669

MRS runs used:

6251 6320 6427 6493 6703  
6256 6321 6428 6630 6705  
6257 6322 6429 6631 6706  
6270 6331 6431 6633 6707  
6271 6346 6432 6636 6713  
6277 6394 6433 6637 6714  
6278 6398 6434 6638 6715  
6279 6399 6435 6639 6722  
6289 6400 6442 6669 6723  
6300 6401 6443 6673 6740  
6302 6405 6444 6674 6741  
6303 6410 6445 6675 6748  
6313 6411 6446 6676 6750  
6314 6415 6447 6677 6751  
6315 6417 6448 6682 6752  
6316 6418 6449 6689 6753  
6318 6419 6475 6690 6754  
6319 6420 6488 6691



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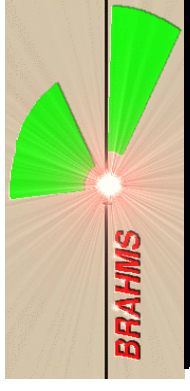


*p+p ratios*  
Bjørn H. Samset

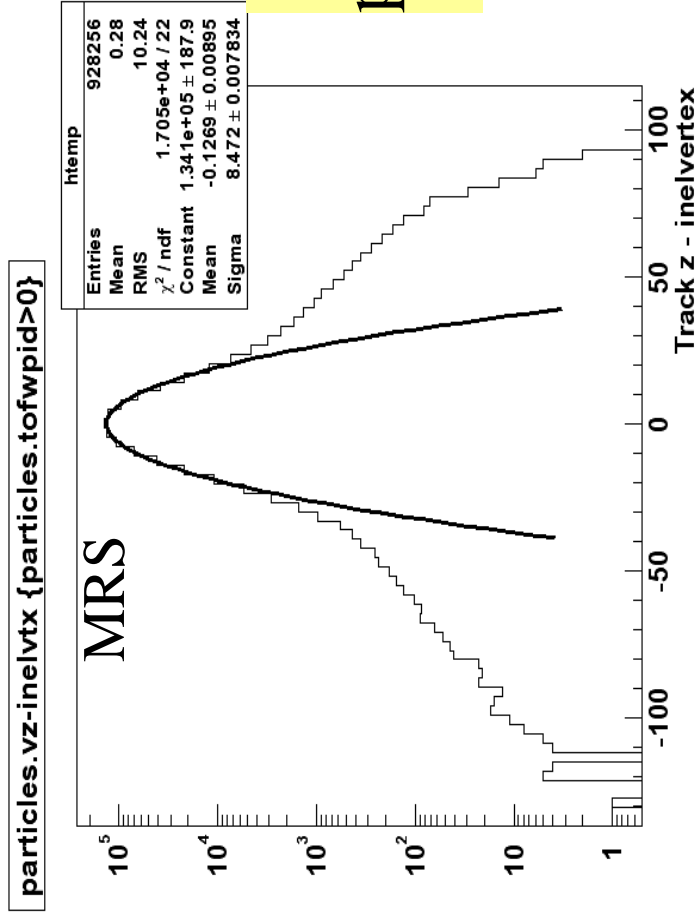
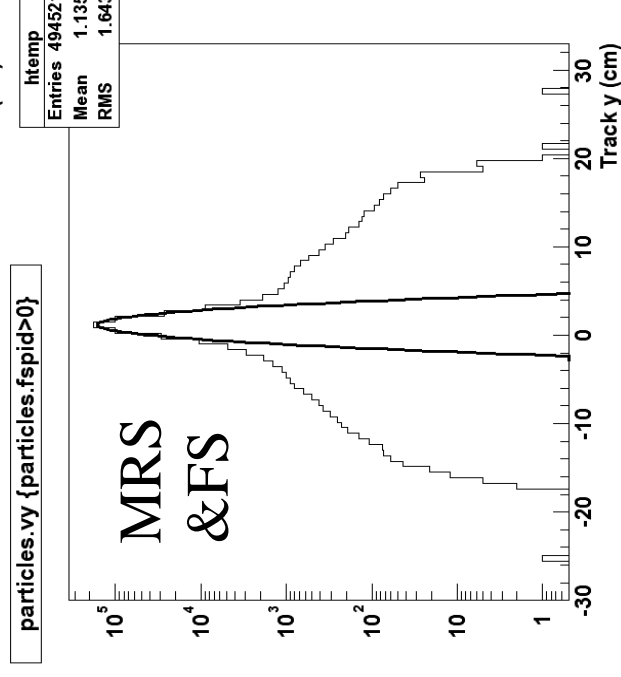
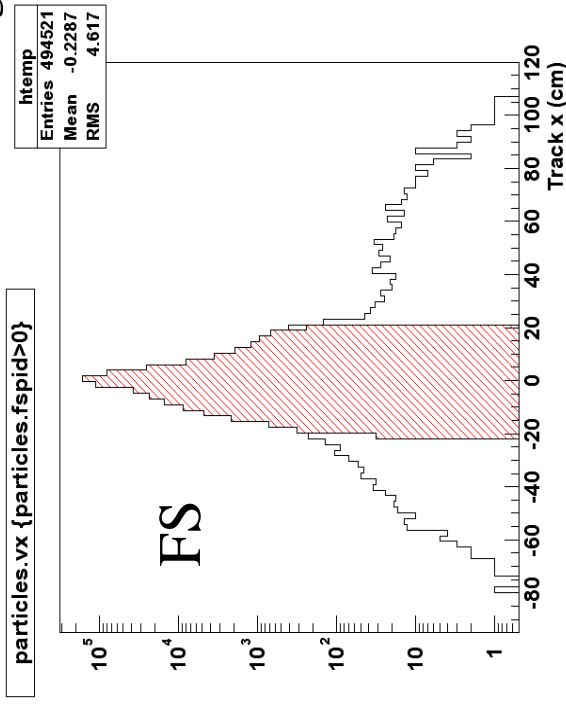


# The analysis

- Select bins in y
- Count the number of IDd particles in each bin, multiply by trigger scaledown (3/6)
- Normalize to the number of NSD triggers (5)
- FS: Divide A field by B field
- MRS:  $r = \sqrt{\text{neg}_A / \text{pos}_B \cdot \text{neg}_B / \text{pos}_A}$

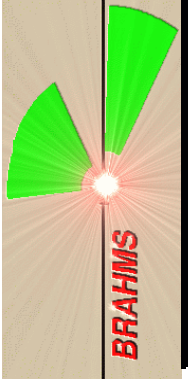
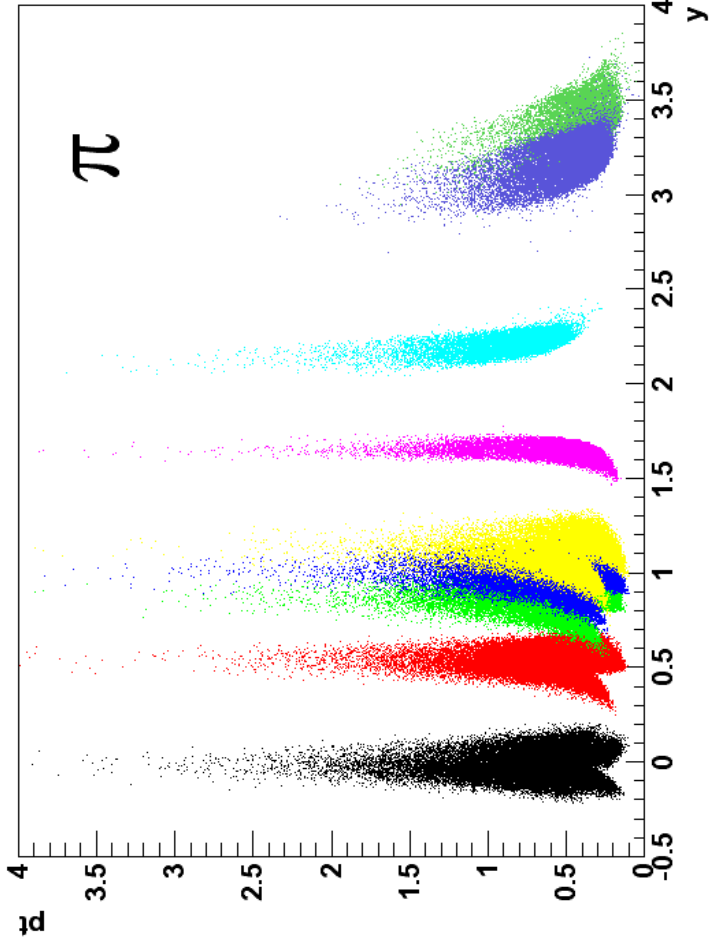


Collaboration meeting



Track projection cuts

# The analysis - $y$ bins

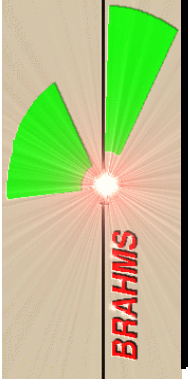


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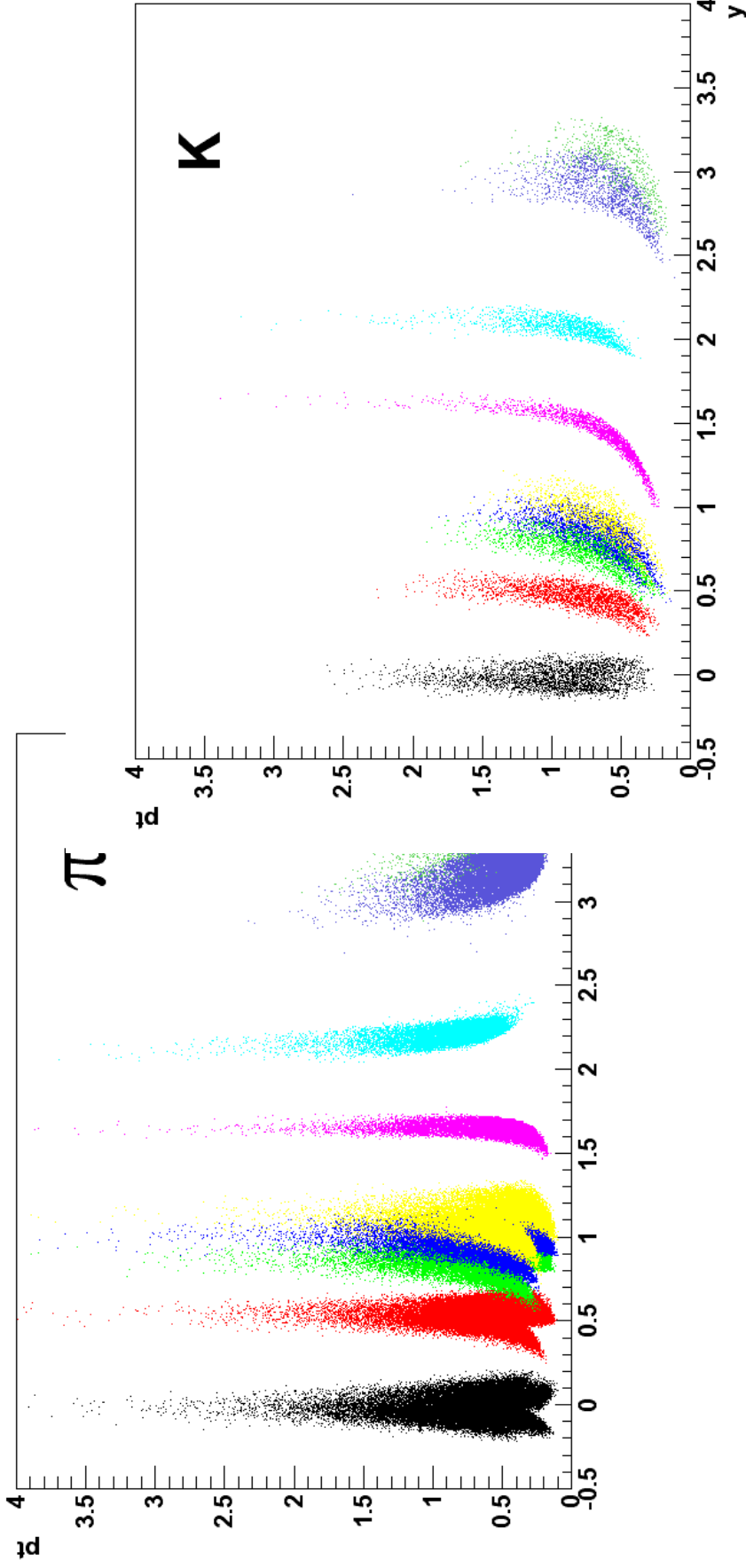


*p+p* ratios  
Bjørn H. Samset

# The analysis - $y$ bins

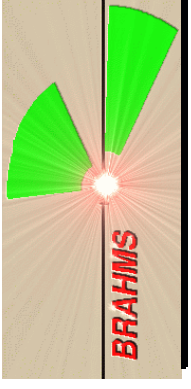


Collaboration meeting  
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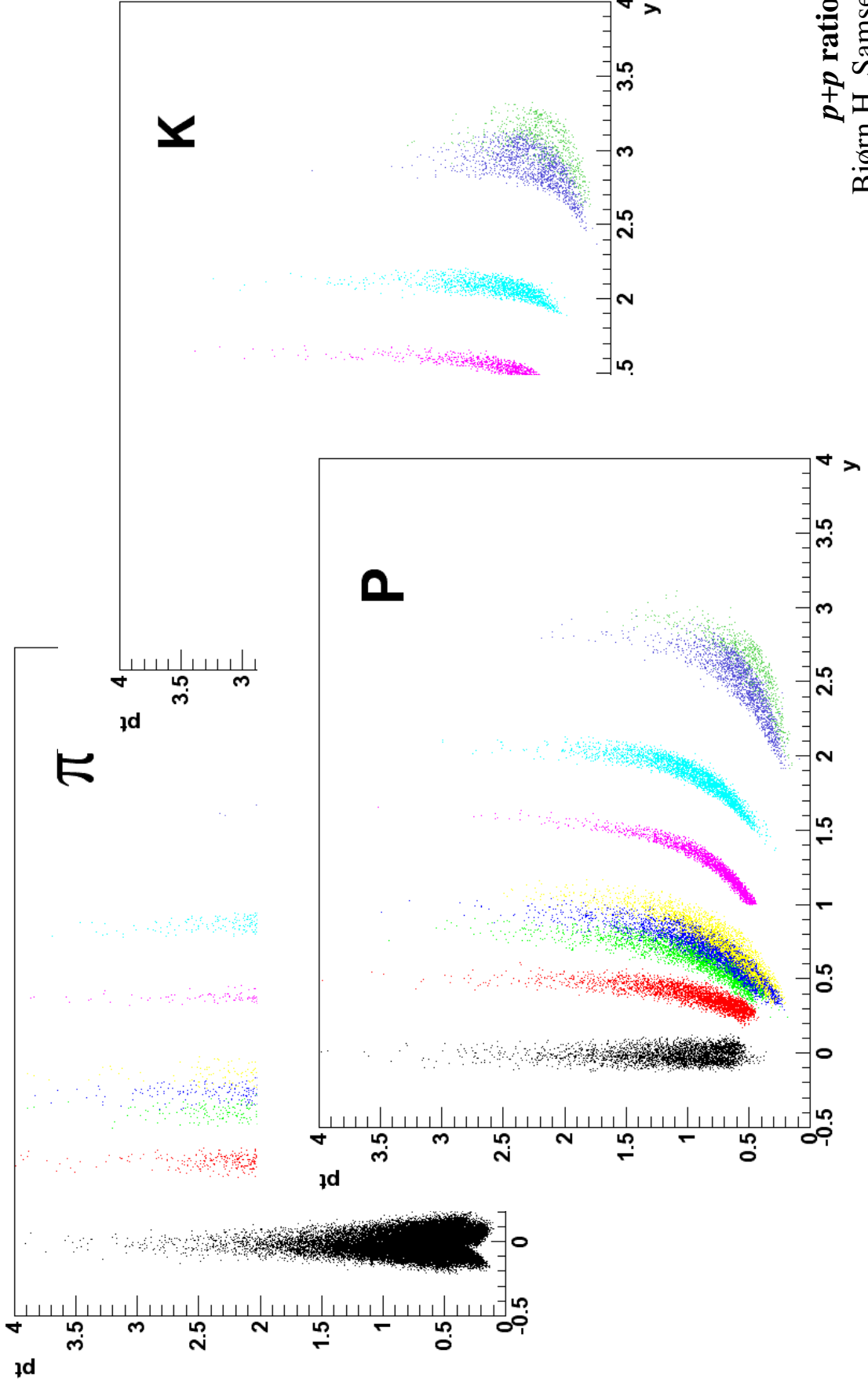


*p+p* ratios  
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# The analysis - $y$ bins



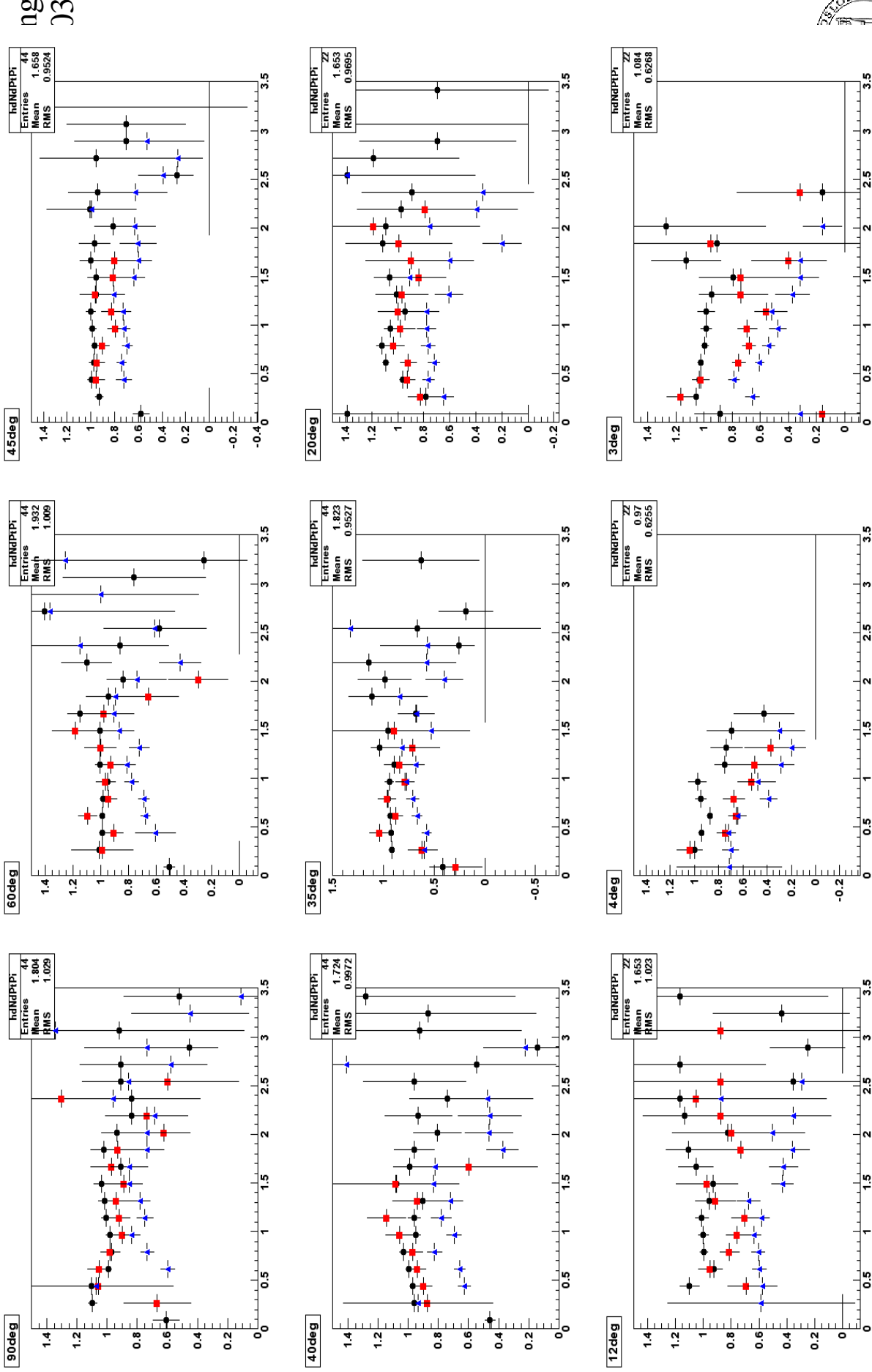
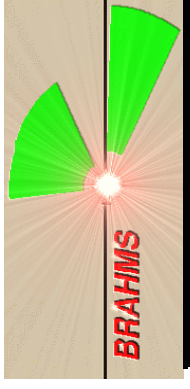
Collaboration meeting  
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*p+p ratios*  
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# pt dependence?

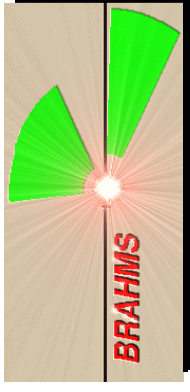
(All plots: Ratio vs. pt)



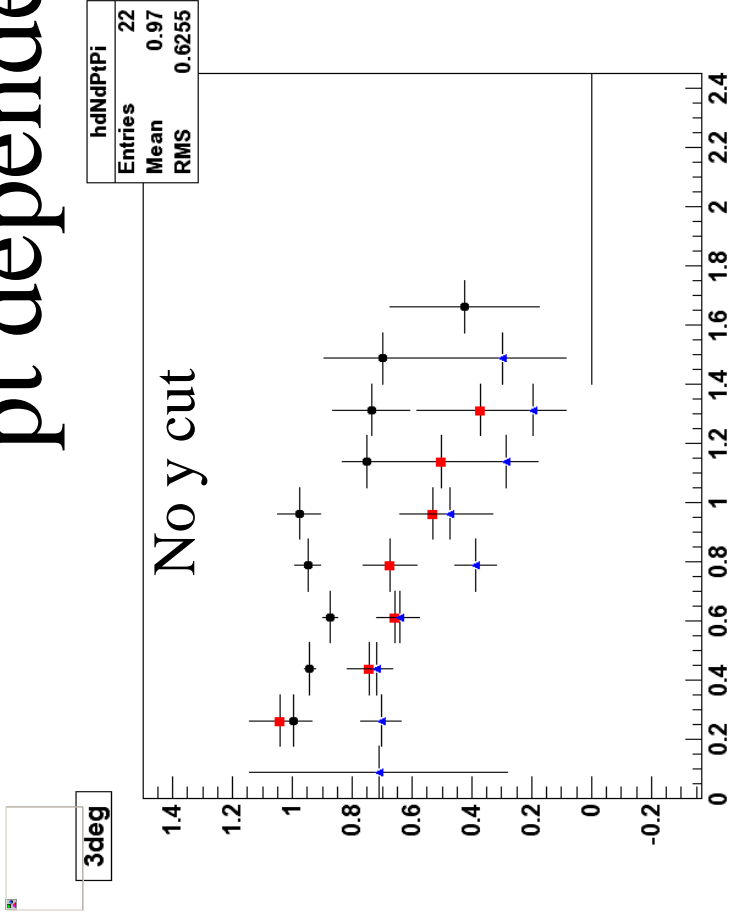
ng  
03



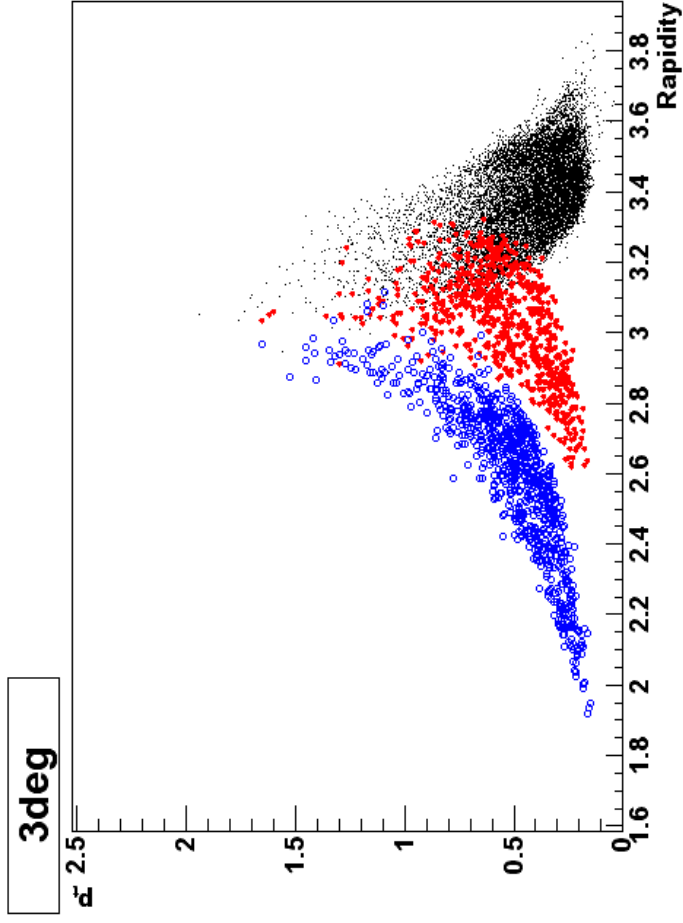
# pt dependence?



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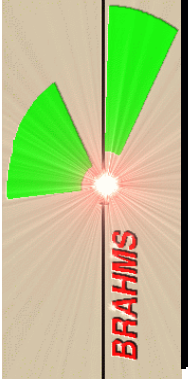
$dy = 0.2$



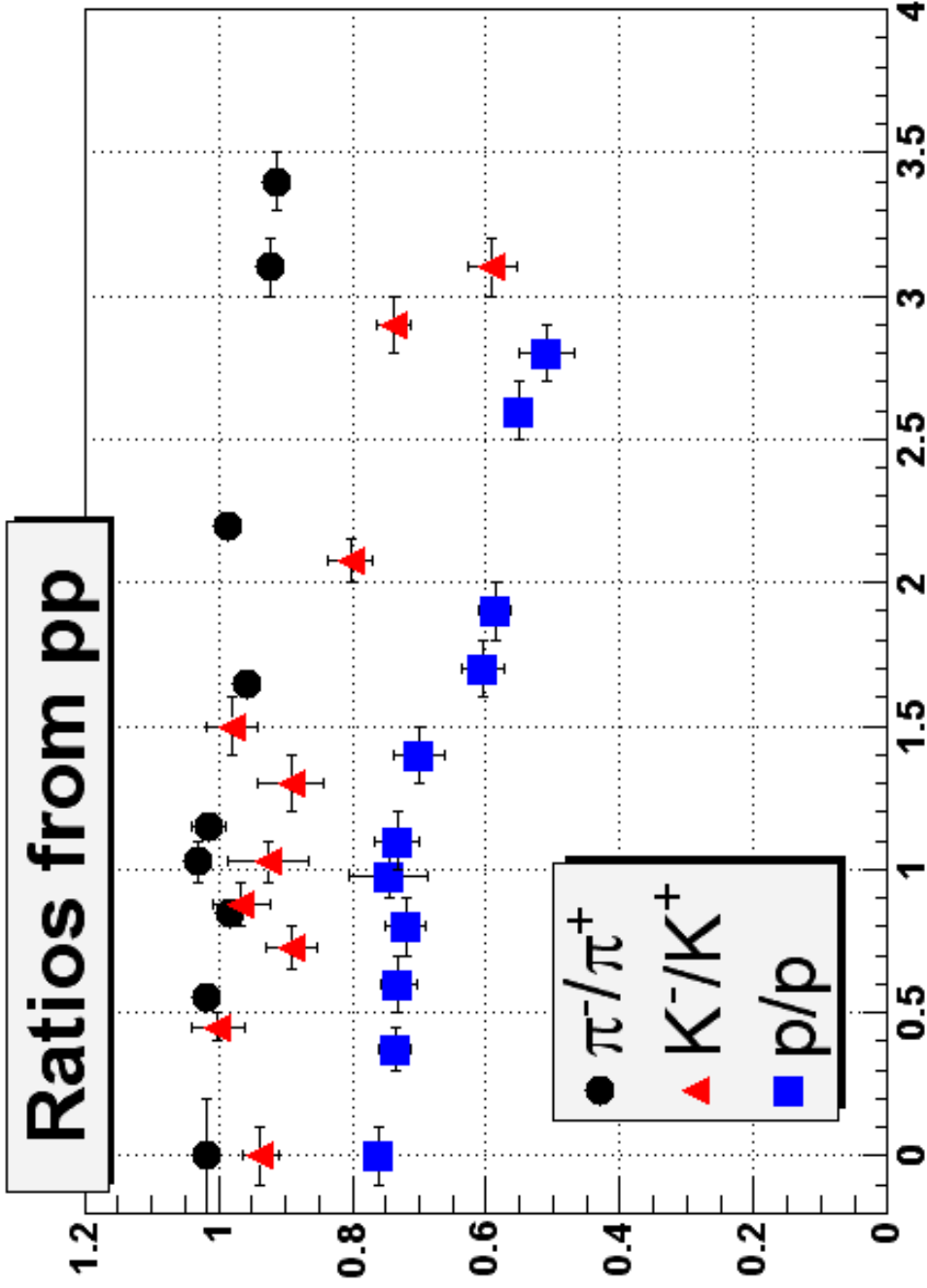
$\pi$ , K: 3.0-3.2  
p: 2.8-3.0

*p+p ratios*  
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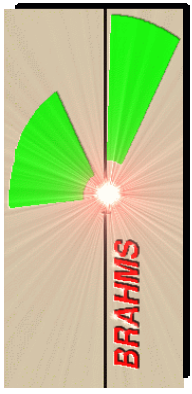
Corrections? None yet.

- Most cancel
- Absorption/production: As for AuAu, except we must add trigger slats.

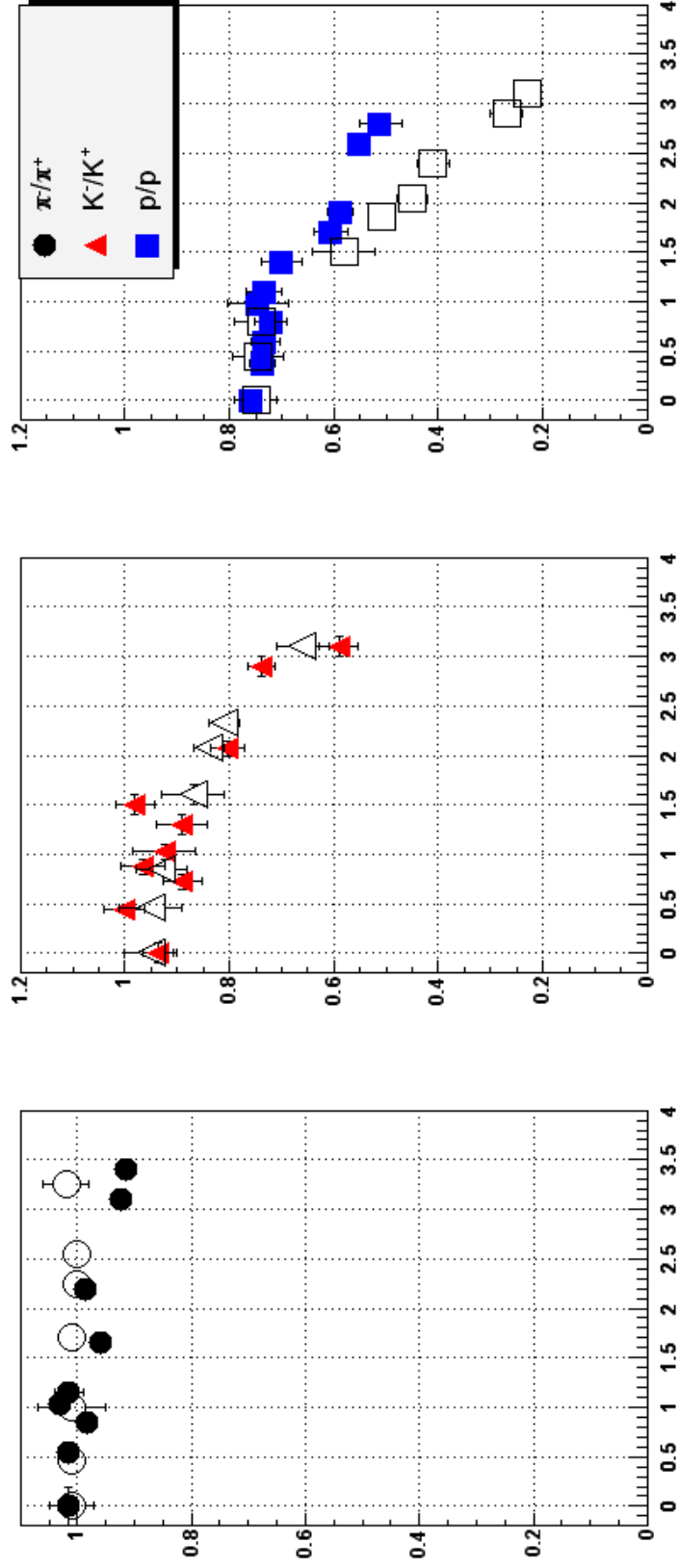


*p+p ratios*  
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# Ratios vs AuAu



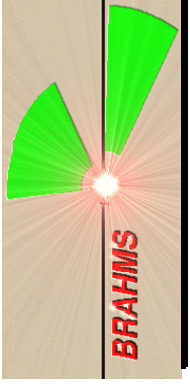
Collaboration meeting  
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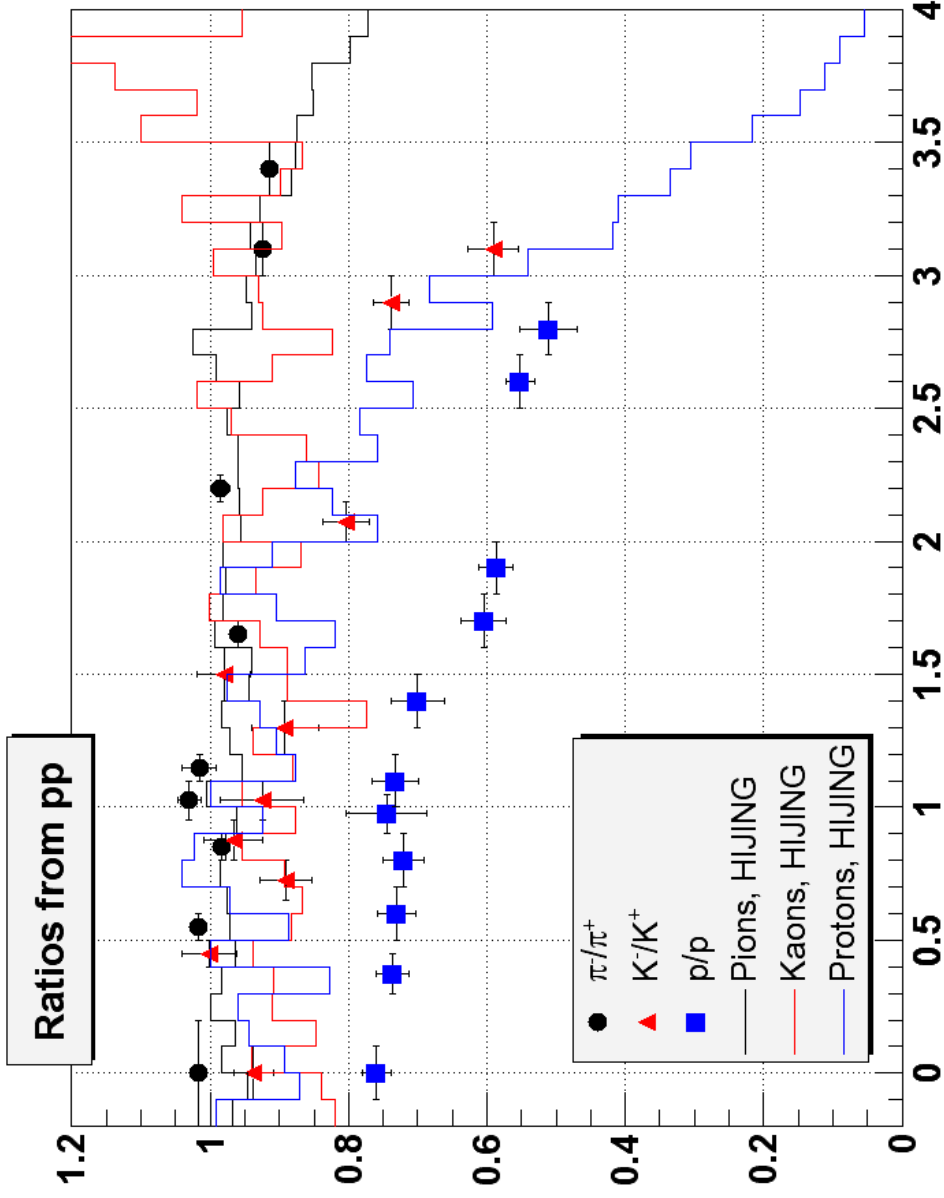
*p+p ratios*  
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# Ratios vs HIJING

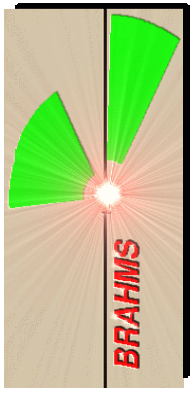


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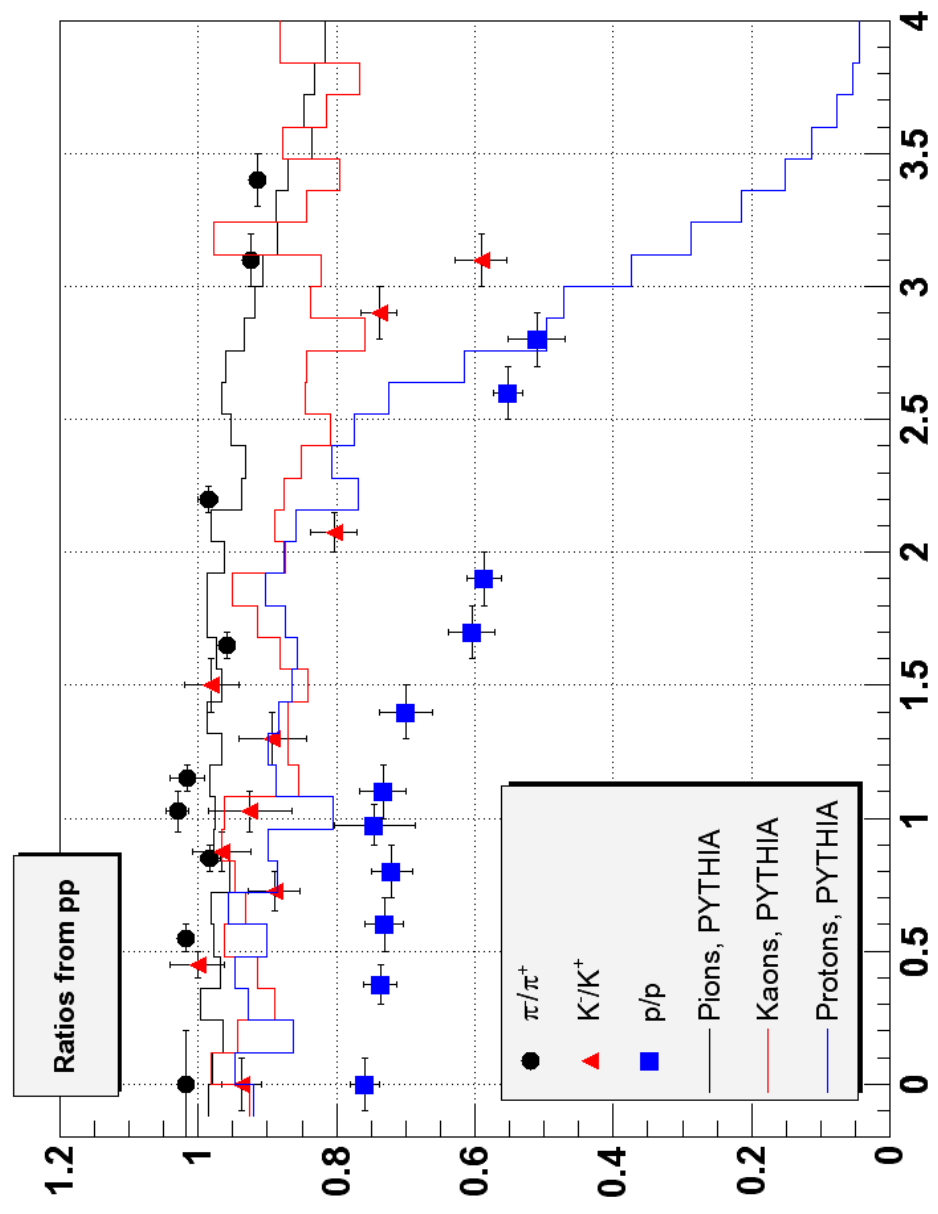


*p+p ratios*  
Bjørn H. Samset

# Ratios vs PYTHIA

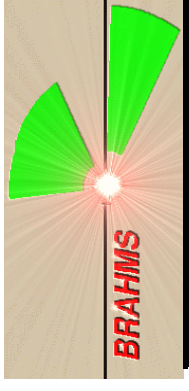


Collaboration meeting  
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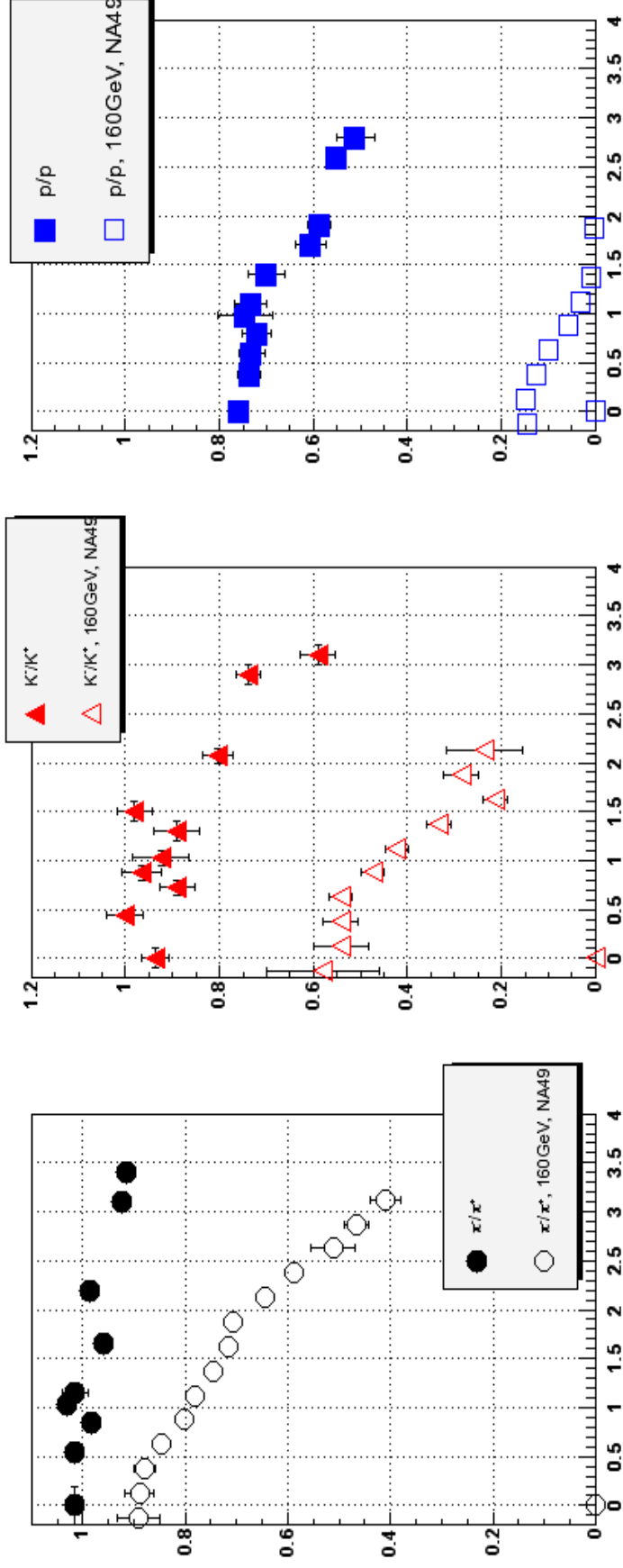


*p+p ratios*  
Bjørn H. Samset

# Ratios in BRAHMS and NA49 (p+p, 160GeV f.t.)



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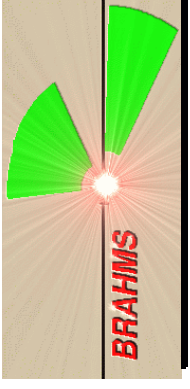
NA49 data:

<http://f.home.cern.ch/f/fsikler/public/ips/near/report/report.ps>

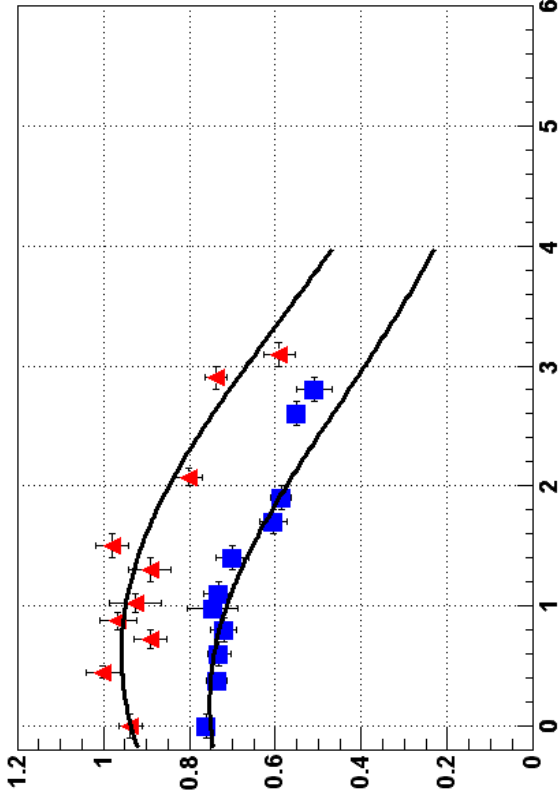


*p+p ratios*  
Bjørn H. Samset

# Chemical plot



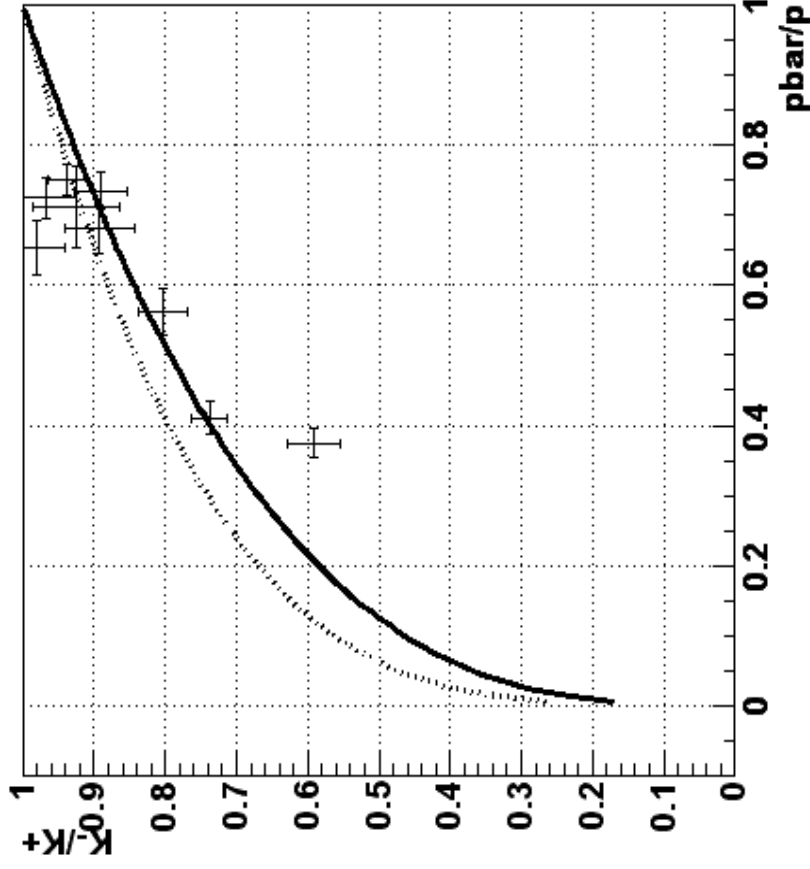
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- The data indicate

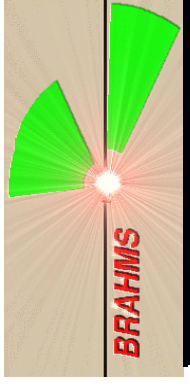
$$K^-/K^+ = (\bar{p}/p)^{1/3}$$

- $\mu_s = 0$ ?
- (Well... Rather shaky as of yet...)



*p+p ratios*  
Bjørn H. Samset

# Issues and cave(b)ats



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The Big Brown Bat is one of seven bat species that utilize caves in Indiana. Bats are found in Most Indiana caves (photo by Sam Frushour).

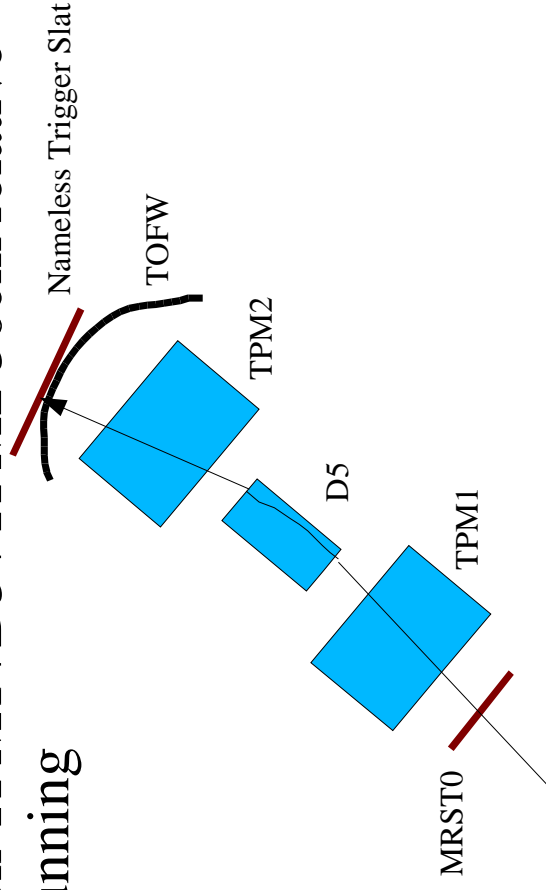
- Trigger efficiency? (What frac. of NSD cross section do we see for trig. 3,5,6?)
- Should add 2003 dataset, especially 8deg data
- Systematic uncertainties
  - › Large (up to 50%) variations between vertex bins for both spectrometers when using inelvertex.
  - › Low-y part of 3deg and 4deg settings gives very high ratios.



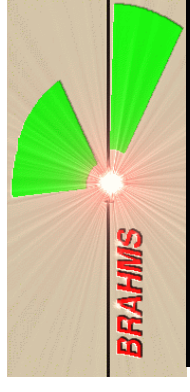
*p+p ratios*  
Bjørn H. Samset

# MRS acceptances

- Using the software in  
brahms\_app/pc\_app/brag/generate  
brahms\_app/pc\_app/brag/acceptance  
with a few modifications:
- Added MRST0 (or was it TMrf?)
  - Cut in TOFW slat to simulate the trigger slat behind the TOFW
  - Pushed back TPM1+D5+TPM2 50cm relative to AuAu running



**Vertex:  
Using track z-  
projection!**

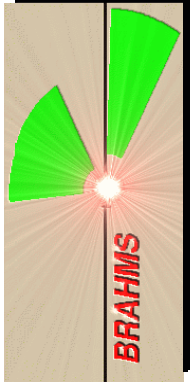


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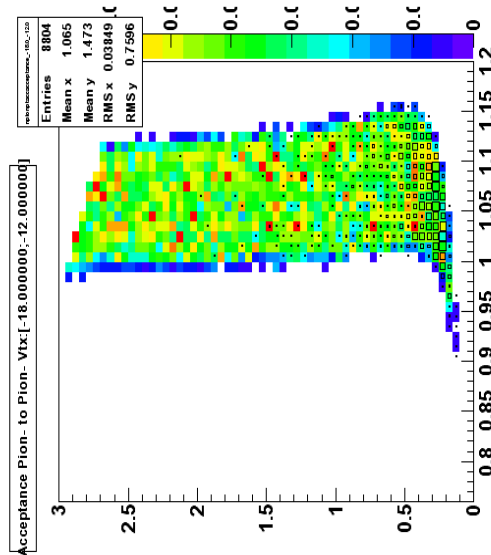
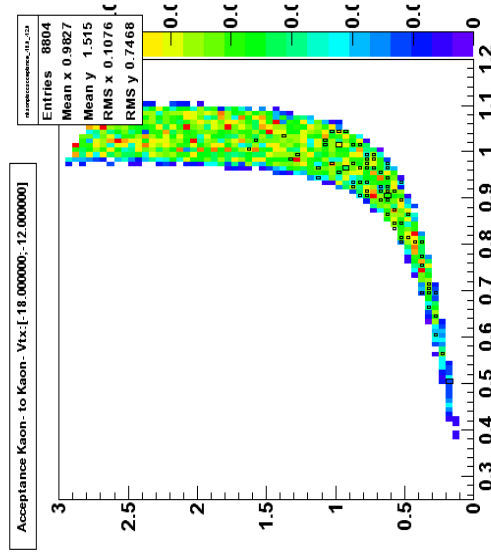
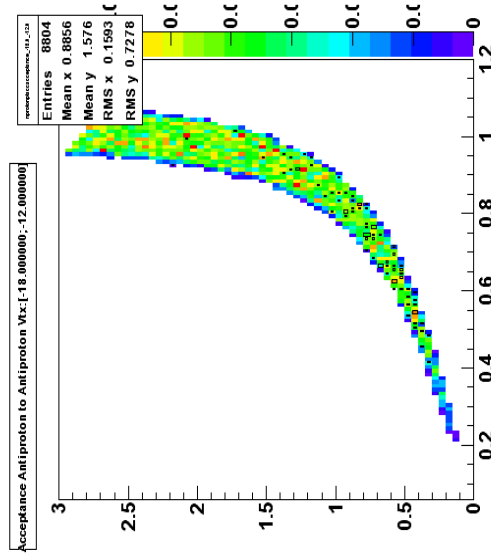
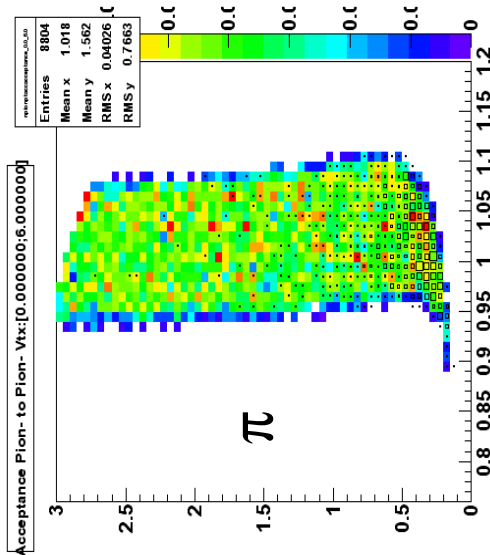
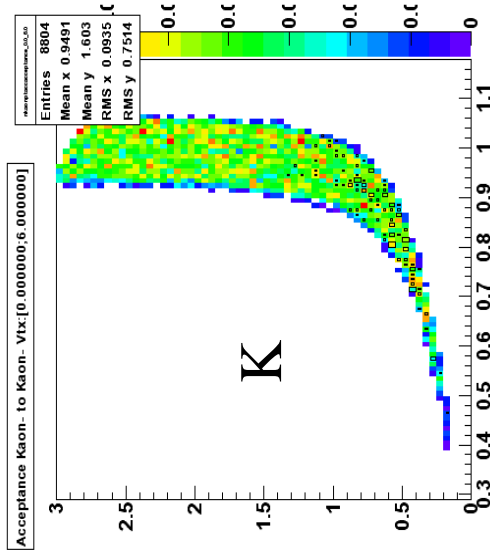
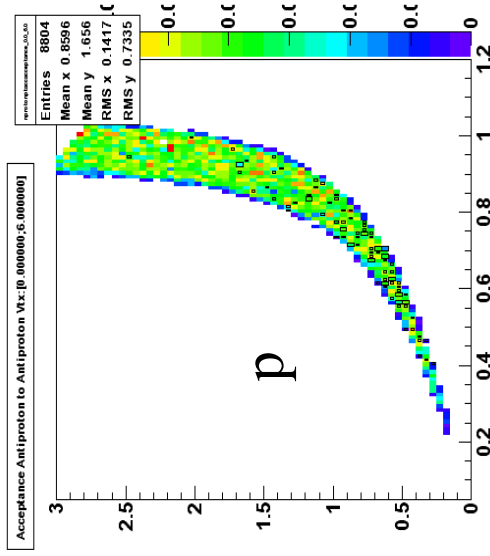


*p+p* ratios  
Bjørn H. Samset

# Acceptance maps - 40deg

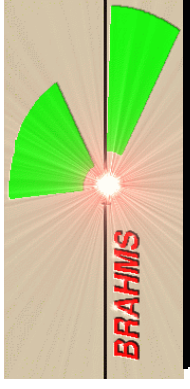


Meeting  
2003

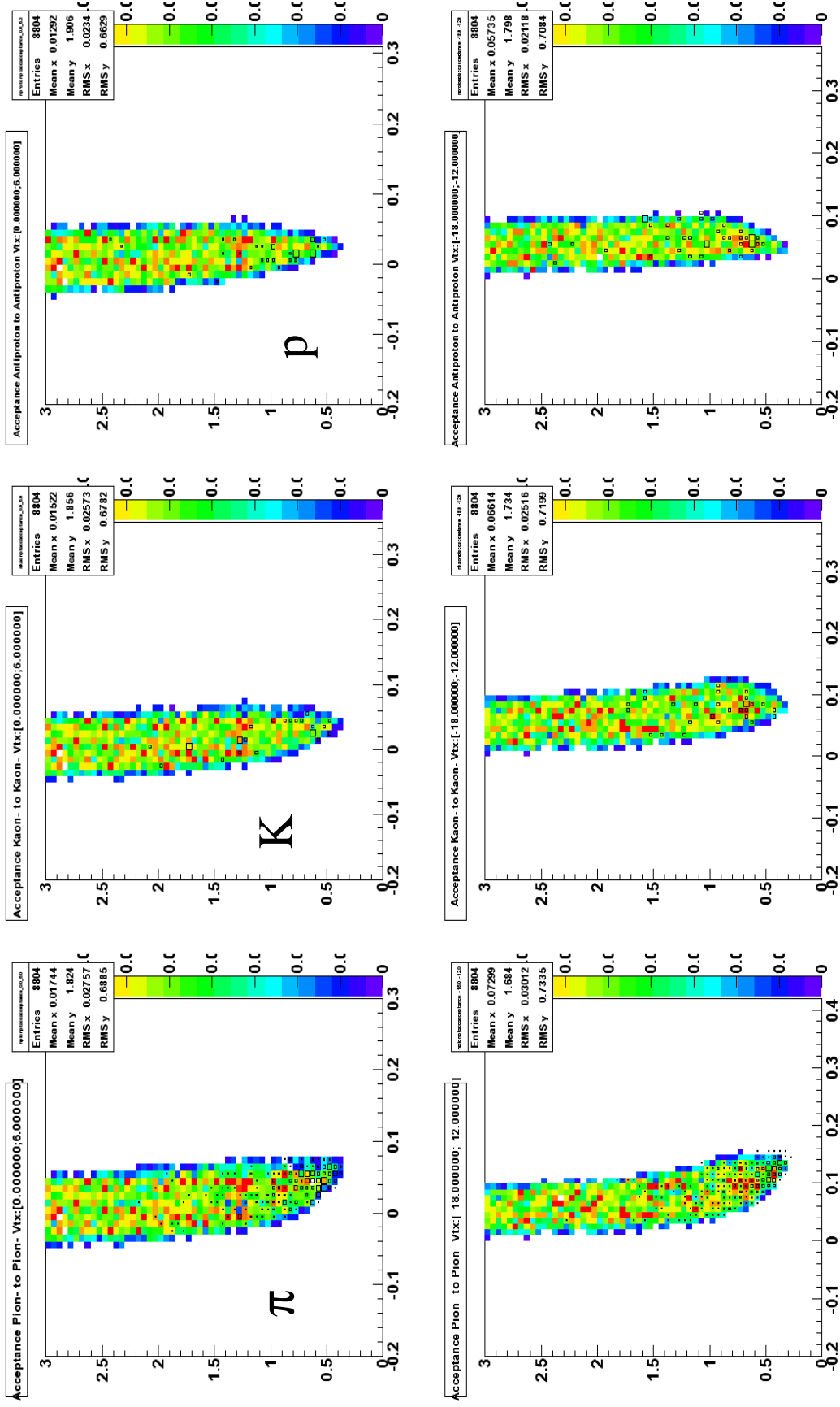


*p+p ratios*  
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# Acceptance maps - 90deg



Meeting  
2003

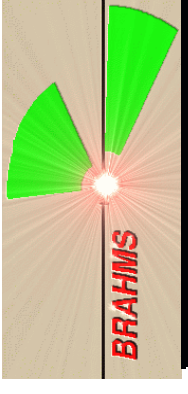


*p+p ratios*  
Bjørn H. Samset



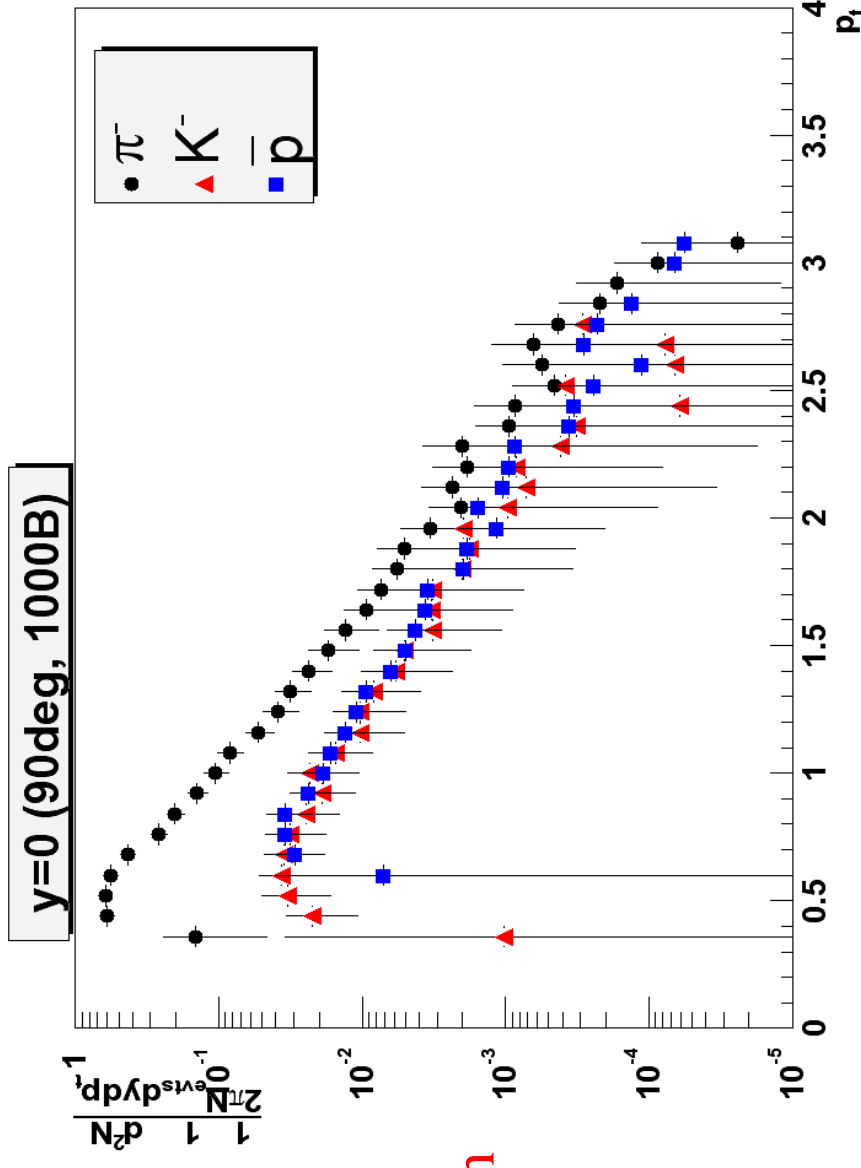
# Spectra at $y=0$ , 1000B setting

(Very recent analysis...)



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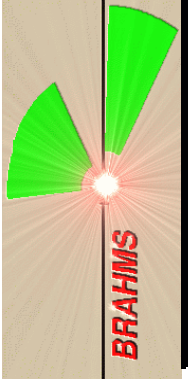
- Acceptance corrections
- Physics corrections  
a'la Eun-Joo
- Cut away very low acc.  
particles etc.
- Effects of trigger slats?
- Acc. file OK? (How high  
was the TOFW trigger  
slat?)



*p+p ratios*  
Bjørn H. Samset

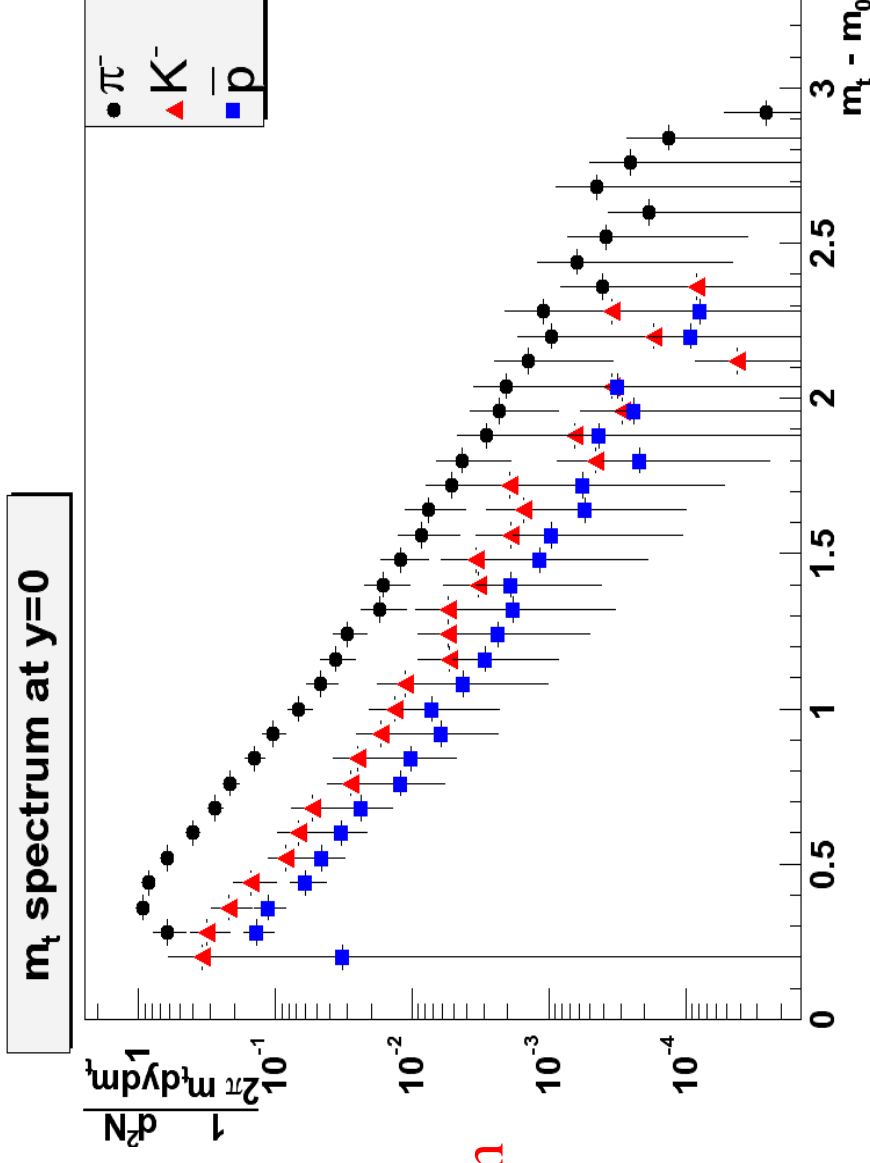
# Spectra at $y=0$ , 1000B setting

(Very recent analysis...)



Collaboration meeting  
Krakow, 4.-7. June 2003

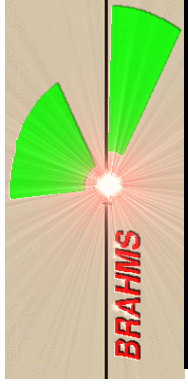
- Acceptance corrections
- Physics corrections  
a'la Eun-Joo
- Cut away very low acc.  
particles etc.
- Effects of trigger slats?
- Acc. file OK? (How high  
was the TOFW trigger  
slat?)



*p+p ratios*  
Bjørn H. Samset

# Spectra at $y=0$ , 1000B setting

(Very recent analysis...)

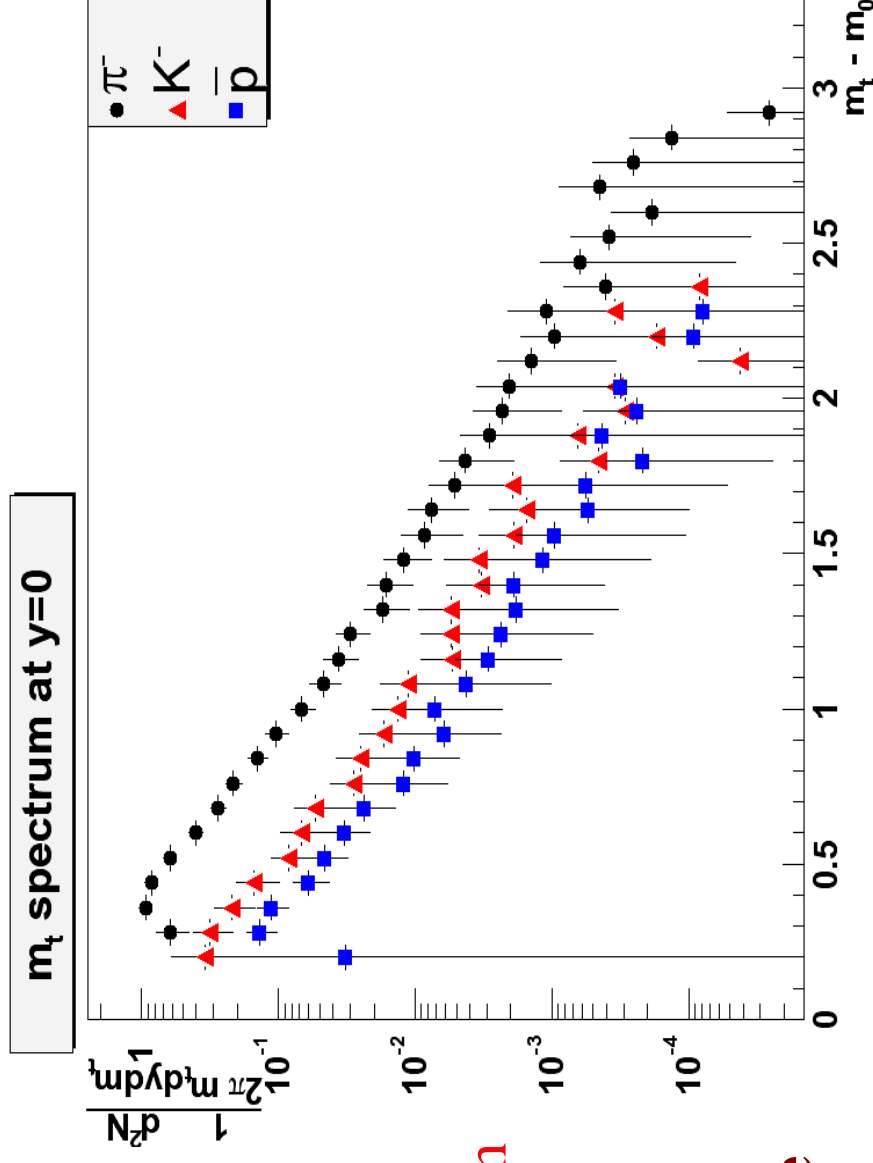


Collaboration meeting  
Krakow, 4.-7. June 2003

- Acceptance corrections
- Physics corrections  
a'la Eun-Joo
- Cut away very low acc.  
particles etc.

- Effects of trigger slats?
- Acc. file OK? (How high  
was the TOFW trigger  
slat?)

**However, the details are  
still wrong. Ratios and  
yields are strange - more work to do!**

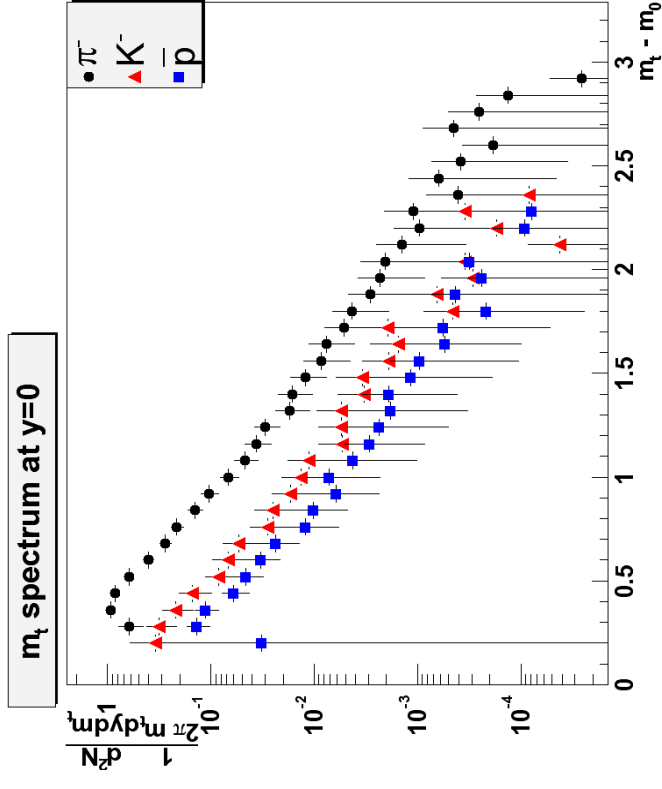
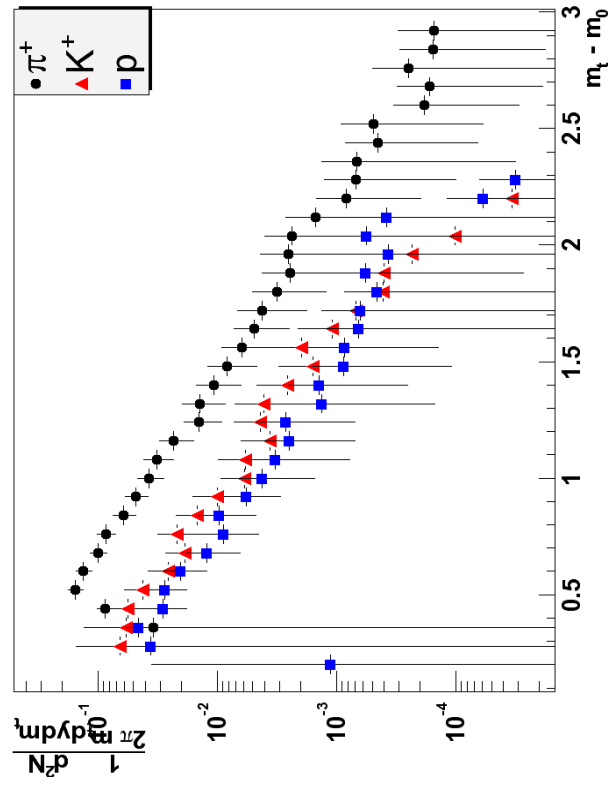
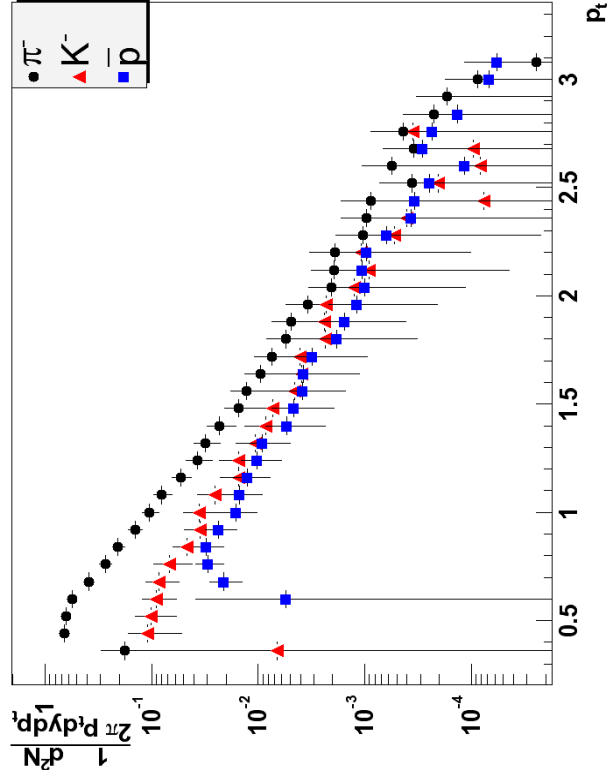
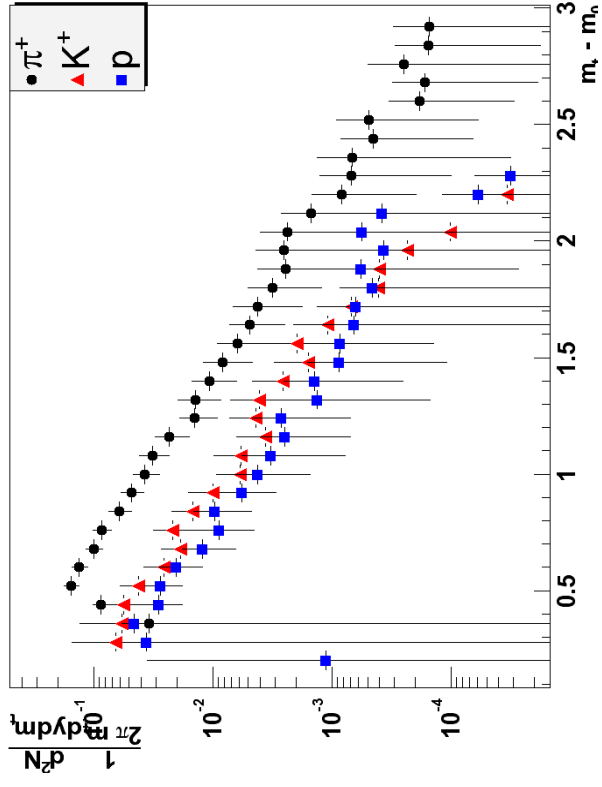


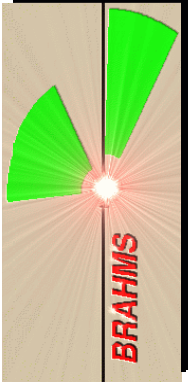
*p+p ratios*  
Bjørn H. Samset

# Spectra at $y=0$ , 1000B setting



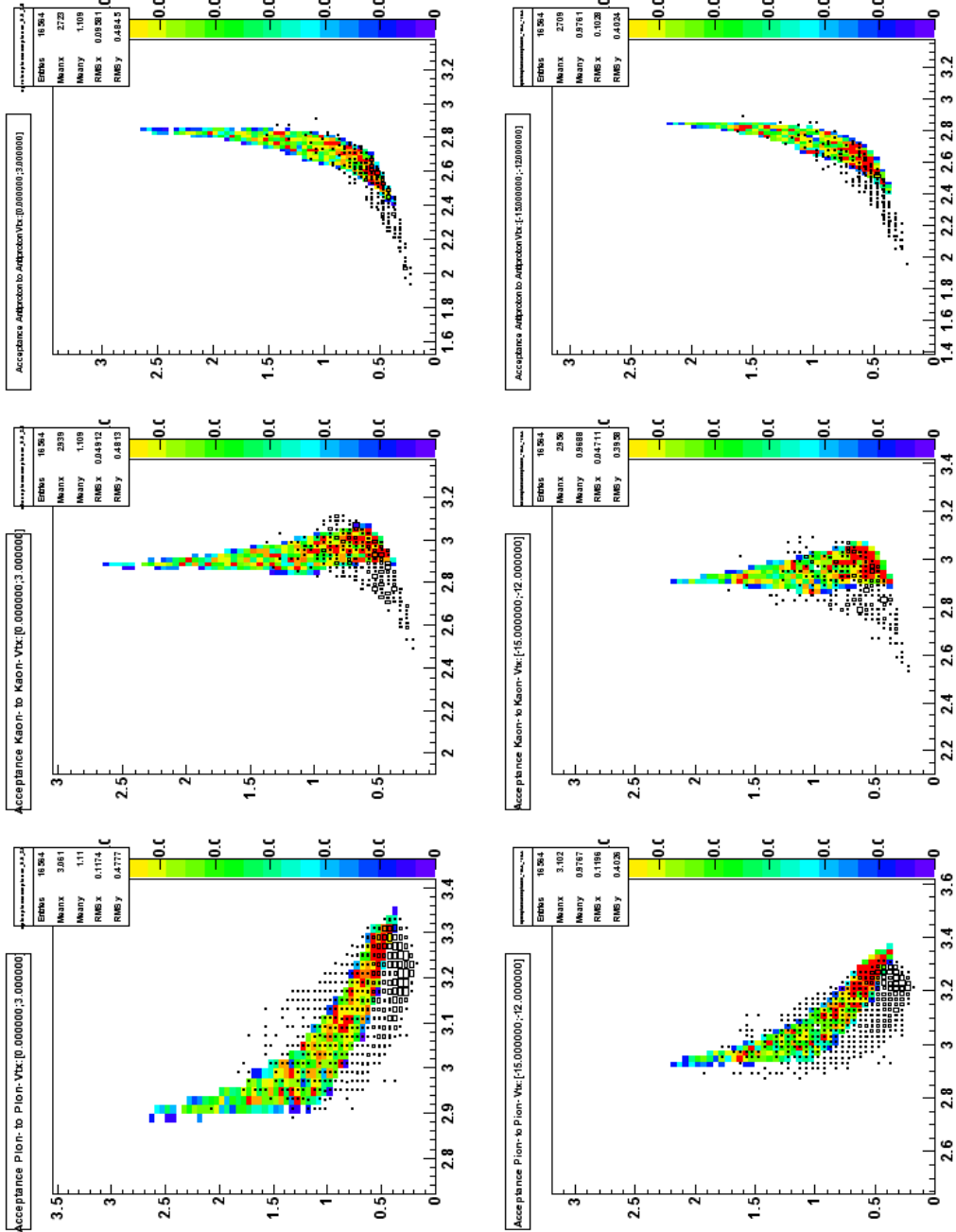
meeting  
10 June 2003





# FS acceptance !?!

Collaboration meeting  
Krakow, 4.-7. June 2003

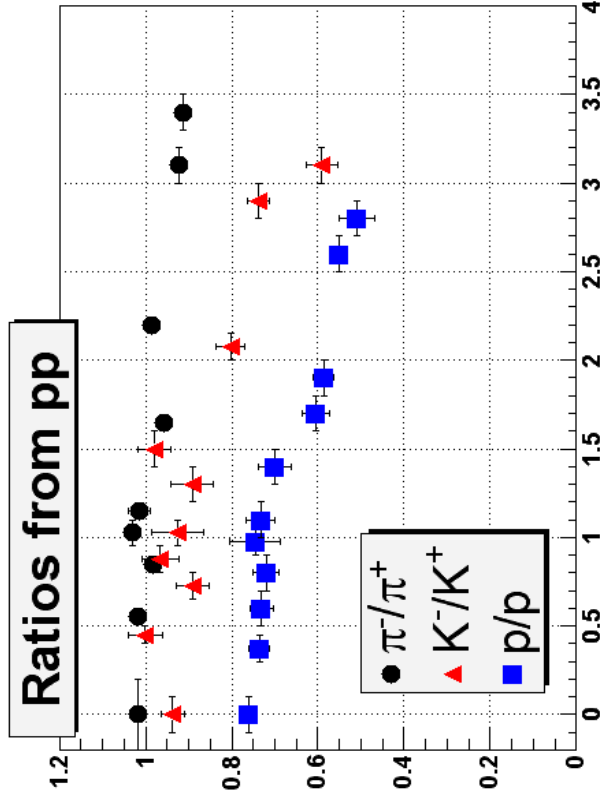


- No track projections here...
- Inelvertex is just too wide
- Acc. based on pos, angle at T1 midplane?
- Works, but how do I get the absolute normalization?

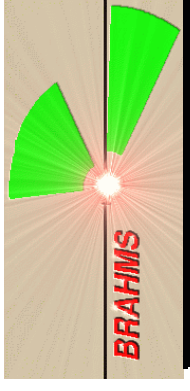


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# Conclusions

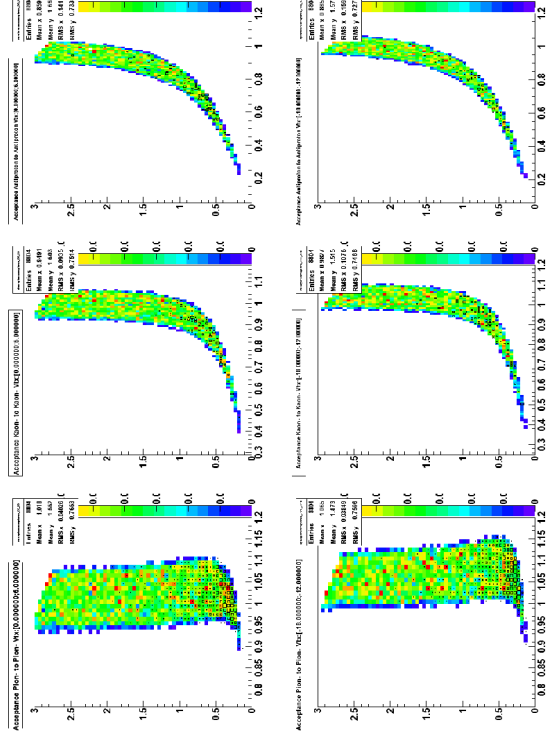


- Ratios are looking good.
- pt dep. at high y?
- Some work still left, but not very much...

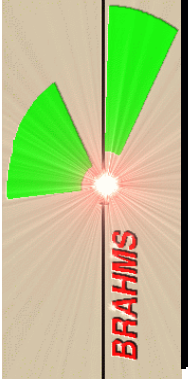


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Krakow, 4.-7. June 2003

- Acceptance is OK for MRS
- Spectra at midrapidity should come soon
- FS is another story...



*p+p* ratios  
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Collaboration meeting  
Krakow, 4.-7. June 2003

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